



Energy Bursts from Fe₈ Molecular Magnet in the Quantum Tunneling Of Magnetization Process

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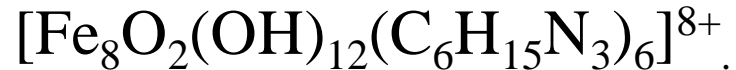
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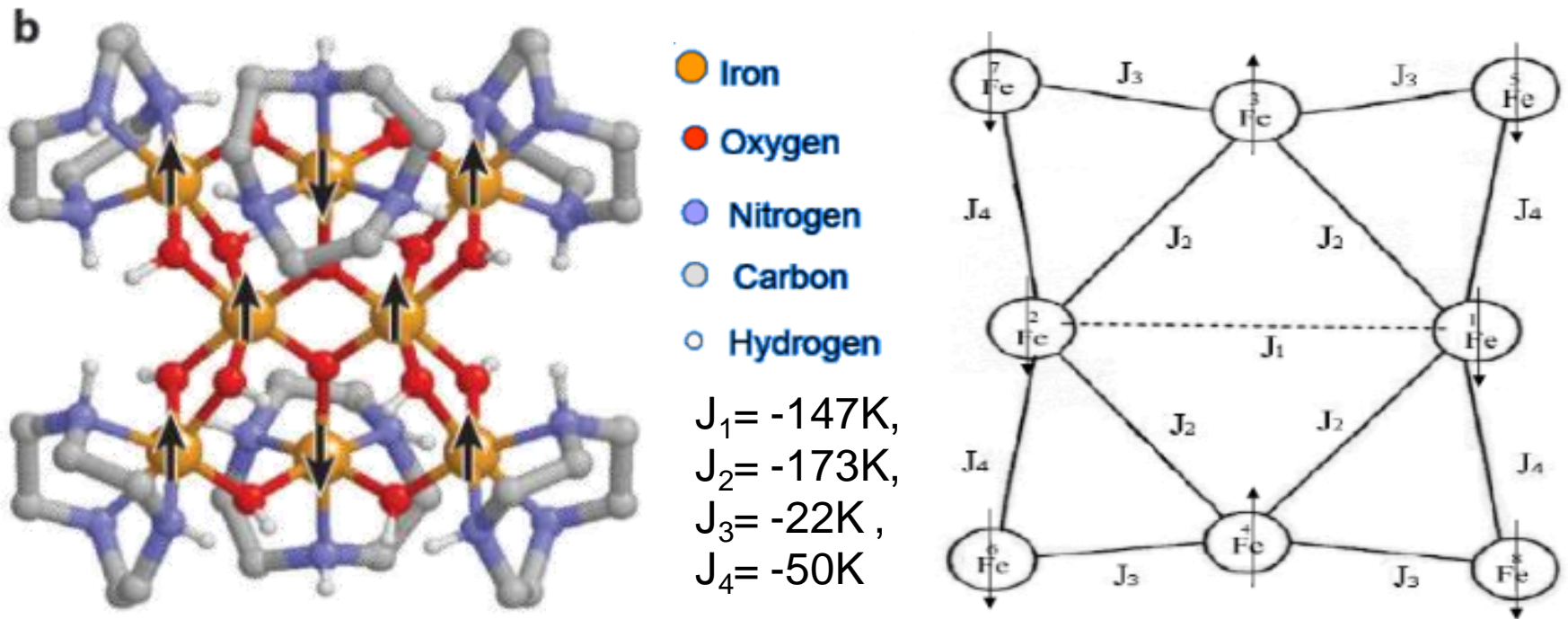
Molecular Magnets

Fe₈ Molecule – Single Molecule Magnet



The antiferromagnetic interactions between the spins of the ions are between 20 to 170K.

Below that the spins are locked

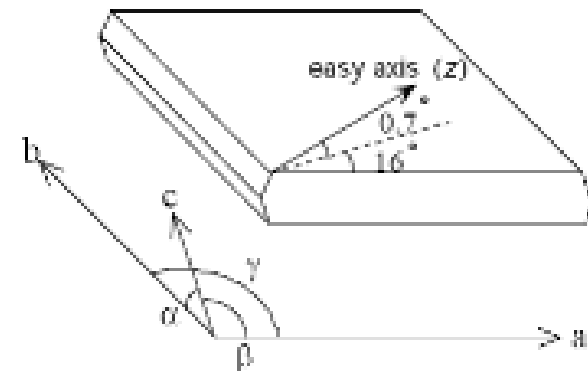
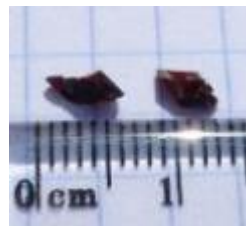


¹K. Wieghardt, K. Pohl, I. Jibril and G. Huttner, *Angew. Chem. Int. Ed. Engl.* 23 (1984), 77.

²C. Delfs,³D. Gatteschi, L. Pardi, R. Sessoli, K. Wieghardt and D. Hanke, *Inorg. Chem.* 32, 3099 (1993).

Fe₈ Molecule – Single Molecule Magnet

- 8 Fe³⁺ ions ($S = 5/2$):
 - 2 ions with spin down: $S = 5$
 - 6 ions with spin up: $S = 15$
 - Total $S = 10$
- Magnetic interactions between the molecules are small ($\sim 0.05\text{K}$).
- Single crystals of Fe₈ are grown from a solvent.
- Orientation is made by a microscope.



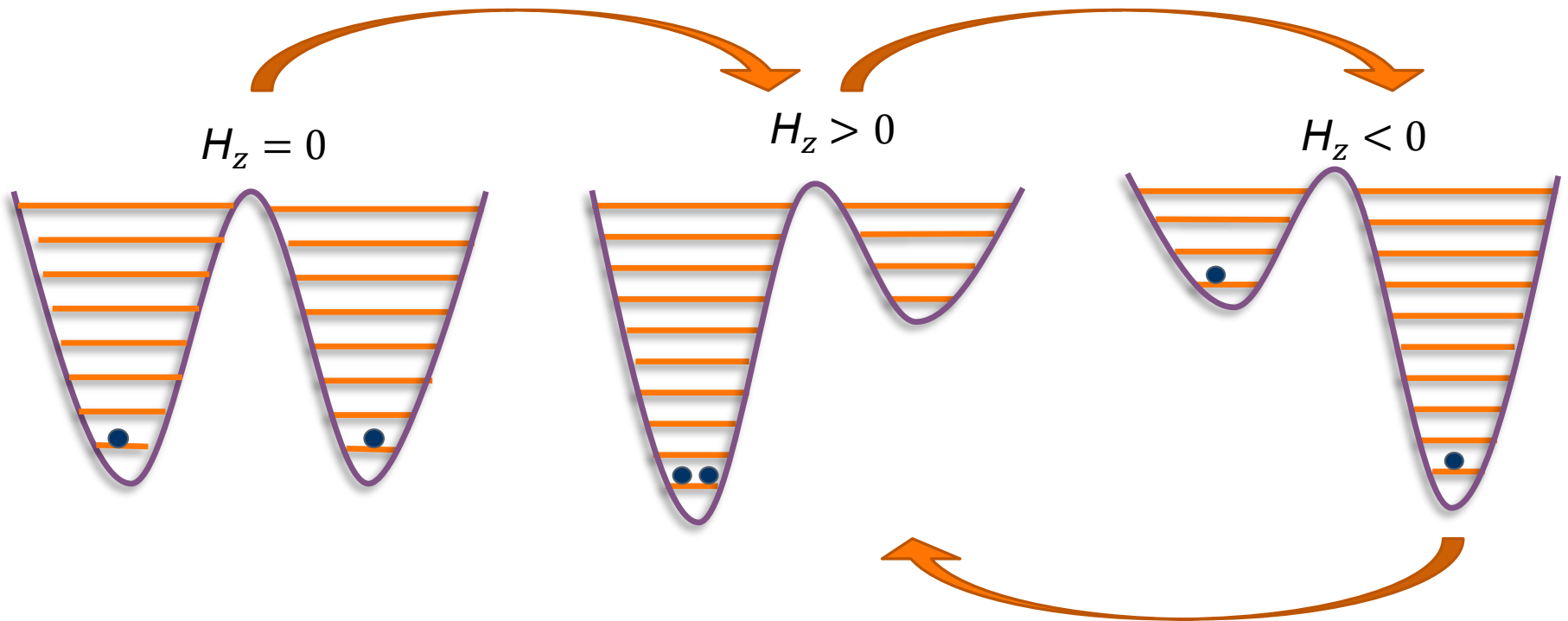
The Crystal field Hamiltonian

$$\mathcal{H} = DS_z^2 + E(S_x^2 - S_y^2) - g\mu_B S_z H_z$$

$D \sim 0.29 \text{ K}$

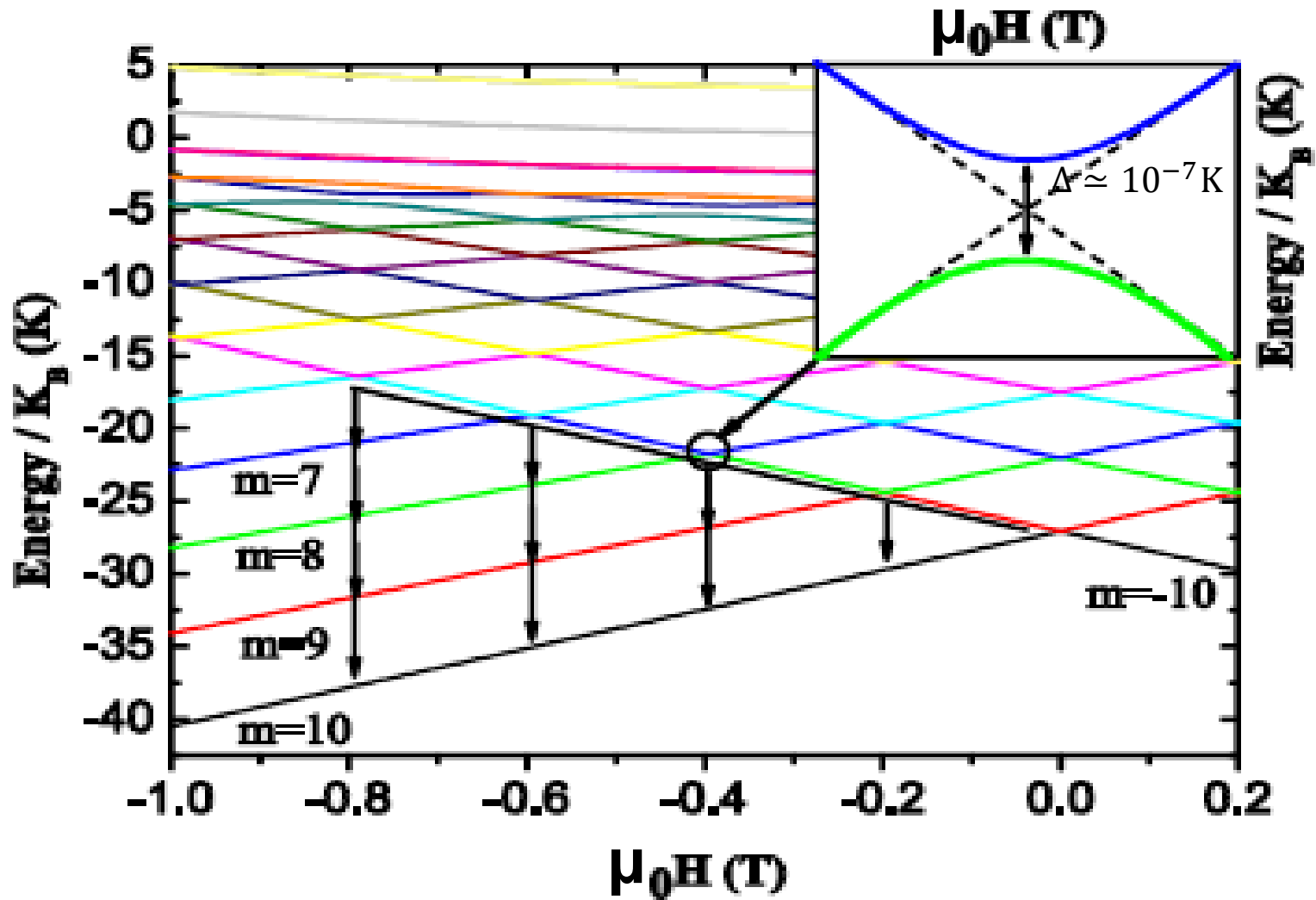
$E \sim 0.045 \text{ K}$

The Experiment



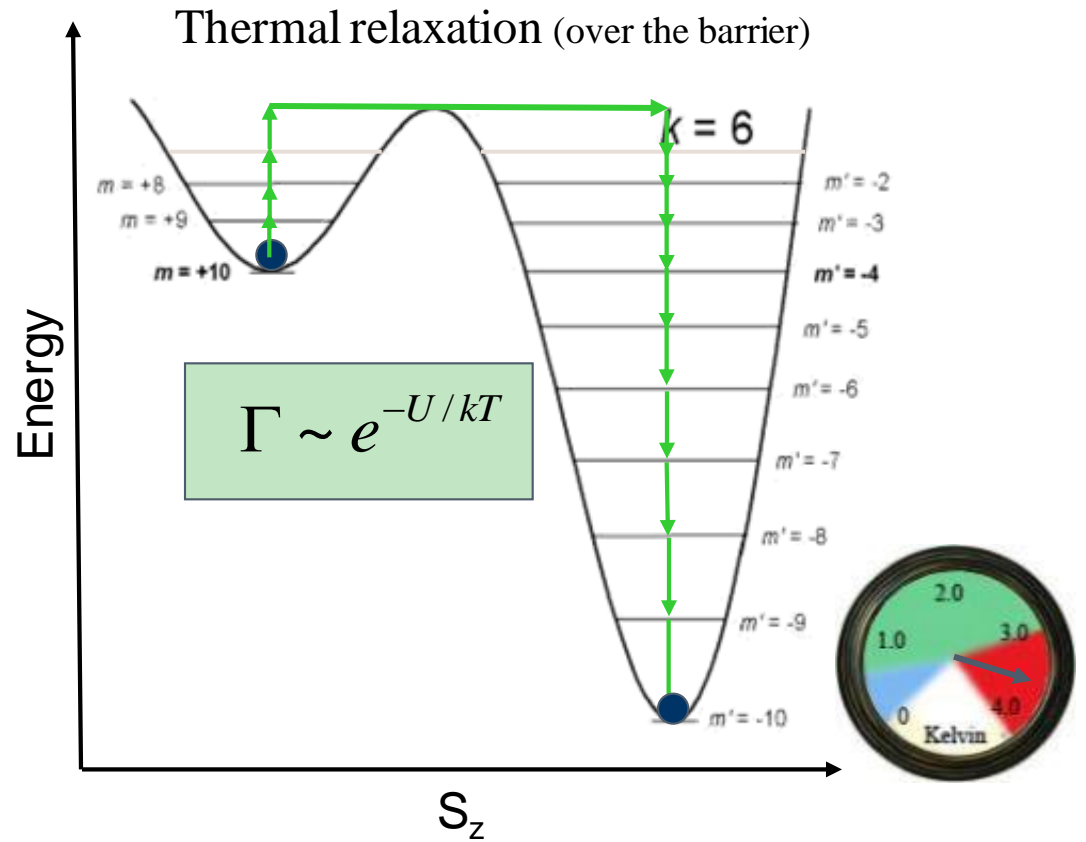
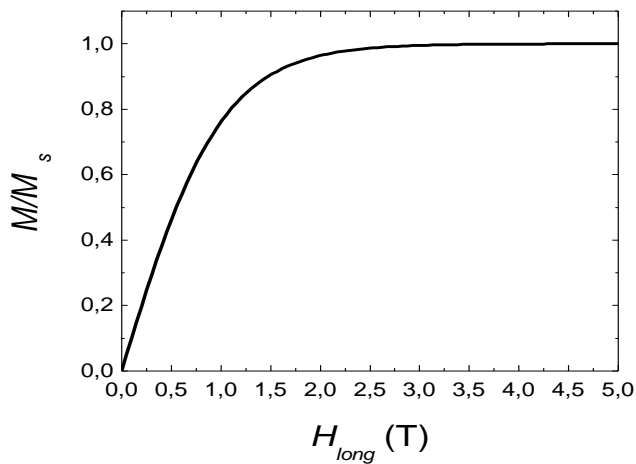
Relaxation processes

Energy Levels for S=10



Relaxation processes

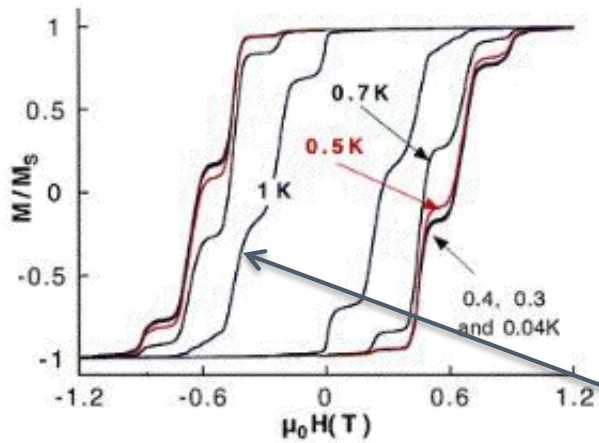
Magnetic field relaxation at high temperature for a single molecule



Relaxation processes

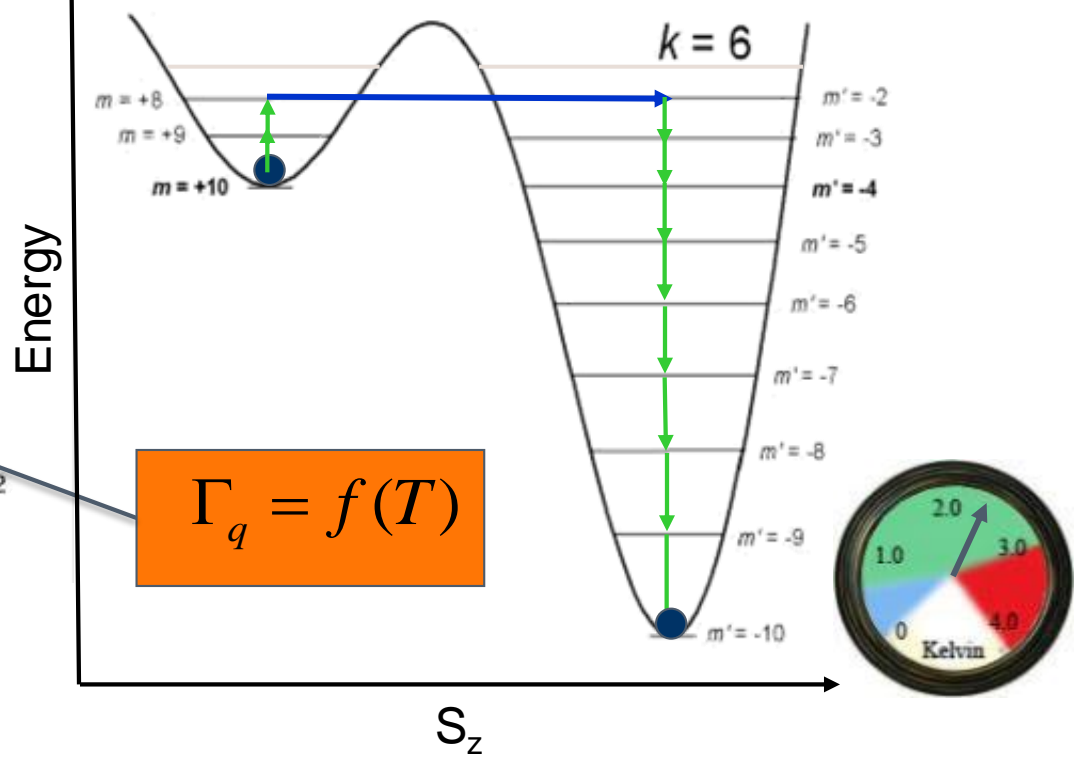
Magnetic field relaxation at low temperature for a single molecule

Hysteresis loop of Fe8



W. Wernsdorfer et al. J. Appl. Phys. 87 (2000), 5481

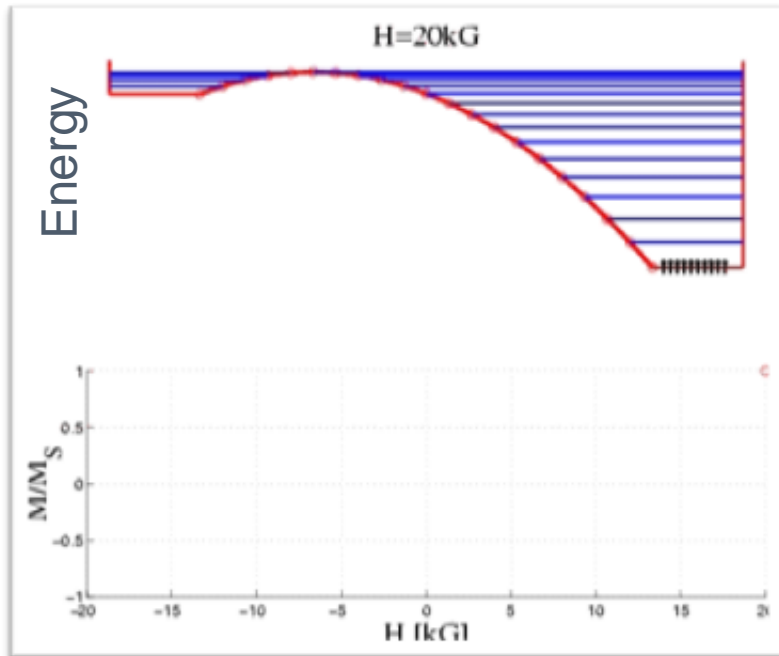
Thermally activated QT (across the barrier)



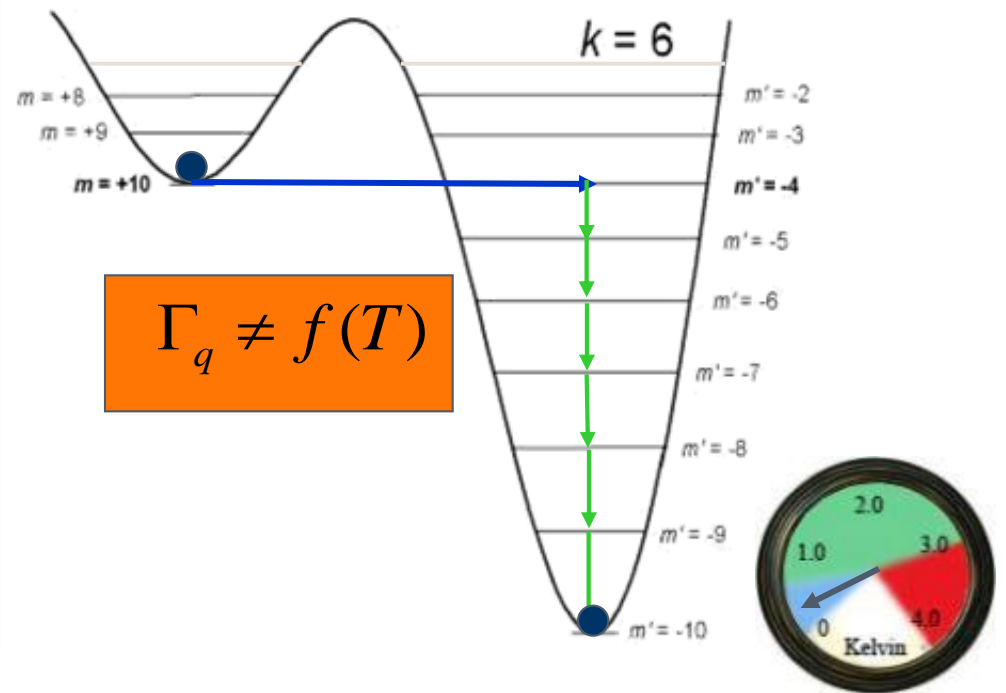
Relaxation processes

Magnetic field relaxation at low temperature for a single molecule

Hysteresis loop of Fe8



Pure Quantum Tunneling (across the barrier)



Relaxation processes

Crystal of Molecular Magnets

At low-temperature magnetic relaxation of whole crystal can occur via two mechanisms:

- A. **Staircase magnetization curve** - Slow mechanism involves quantum transitions at unrelated spatial points.
- B. **Magnetic Avalanche** - Abrupt reversal of the magnetization.

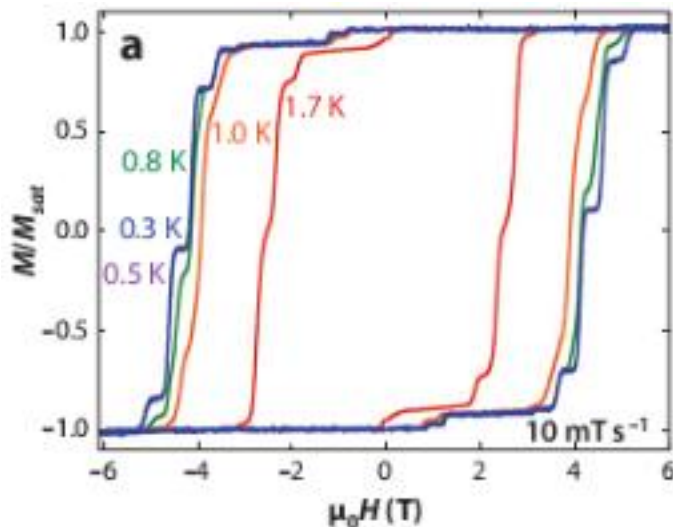


Figure A

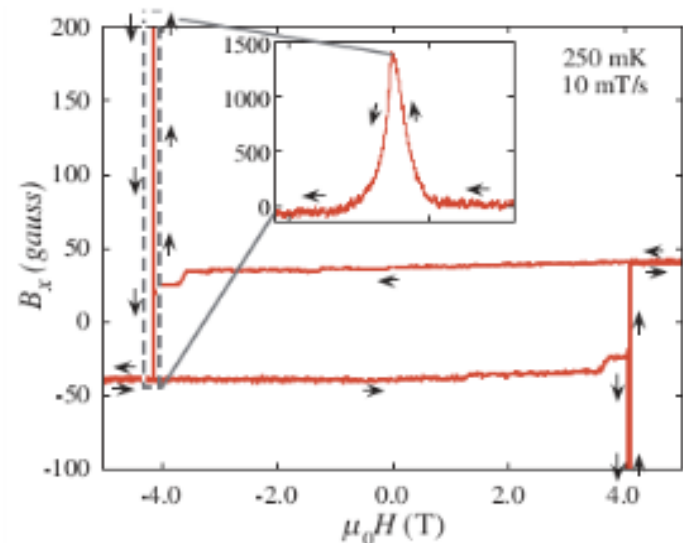
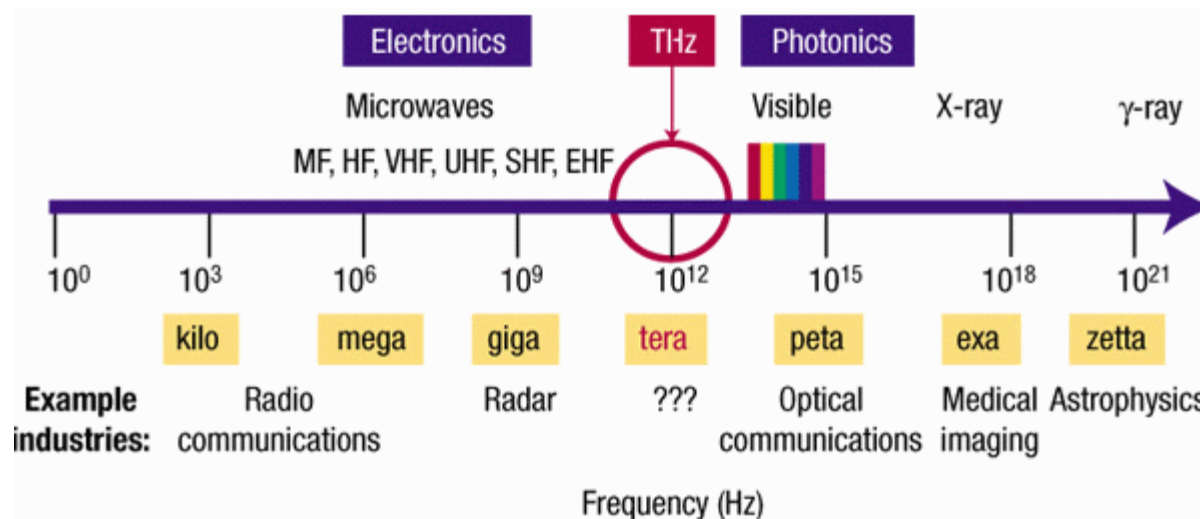


Figure B

Radiation Spectrum

- The expected energy release after the tunneling is twice $\approx 5K \approx 109.5\text{GHz}$ with photon wavelength of about 2.7mm.
 - Cosmic Background radiation is at 160.2 GHz.
 - At frequencies within this range, useful power generation and receiver technologies are inefficient and impractical.



- Our sample size is also about $3 \times 3 \times 1 \text{mm}$ which is the elementary condition for **Dicke Superradiance**.

Plan A : Is There Radiation emanating from Fe₈ ?

VOLUME 89, NUMBER 15

PHYSICAL REVIEW LETTERS

7 OCTOBER 2002

Superradiance from Crystals of Molecular Nanomagnets

E. M. Chudnovsky¹ and D. A. Garanin²

APPLIED PHYSICS LETTERS

VOLUME 84, NUMBER 13

Electromagnetic radiation produced by avalanches in the magnetization reversal of Mn₁₂-acetate

J. Tejada^{a)}

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PHYSICAL REVIEW B **79**, 180404(R) (2009)

Electromagnetic radiation emanating from the molecular nanomagnet Fe₈

Oren Shafir and Amit Keren

Coherence in Spontaneous Radiation Processes

R. H. DICKE

Palmer Physical Laboratory, Princeton University, Princeton, New Jersey

(Received August 25, 1953)

(Enhanced spontaneous emission)

- S&S-radiance arises from spontaneous organization and phase-locking between initially independent systems.
- The effect occurs when the molecules are confined to a volume smaller than radiation wavelength cubed.
- The spontaneous emission is enhanced when the molecules' dipole moments are in phase (super-radiance) and inhibited when they are out of phase (sub-radiance).
- Both effects are universal and fundamental, allowing SR phenomena to occur in a variety of physical systems.

$$\Gamma \simeq \frac{N^2}{4} \Gamma_0$$

Plan B: Are there Avalanches in Fe8?

Avalanche in Mn₁₂

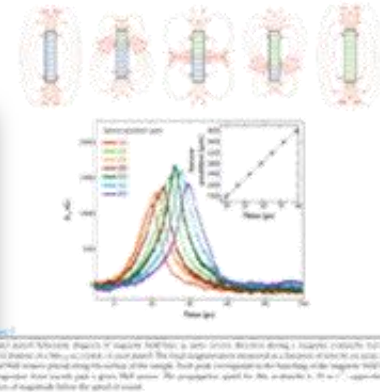
PRL 95, 147201 (2005)

PHYSICAL REVIEW LETTERS

week ending
30 SEPTEMBER 2005

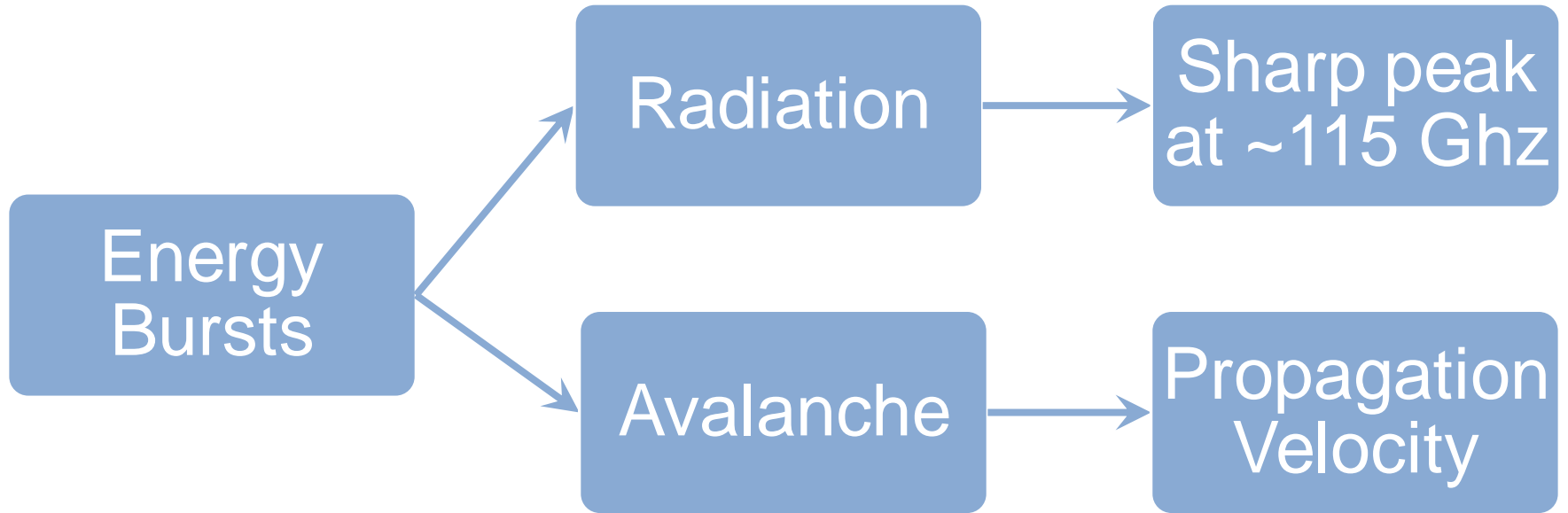
Propagation of Avalanches in Mn₁₂-Acetate: Magnetic Deflagration

Yoko Suzuki,¹ M. P. Sarachik,¹ E. M. Chudnovsky,² S. McHugh,¹ R. Gonzalez-Rubio,¹ Nurit Avraham,³ Y. Myasoedov,³
E. Zeldov,³ H. Shtrikman,³ N. E. Chakov,⁴ and G. Christou⁴



- The Hysteresis loop is temperature depended down to 60 mK
- Avalanches occur in a stochastic way both at resonant magnetic fields (where energy levels on opposite side of the barrier match), and away from resonance.
- The avalanche propagates through the crystal at a constant velocity that is roughly two orders of magnitude smaller than the speed of sound ≈ 10 (m/s).

Experimental Motivation

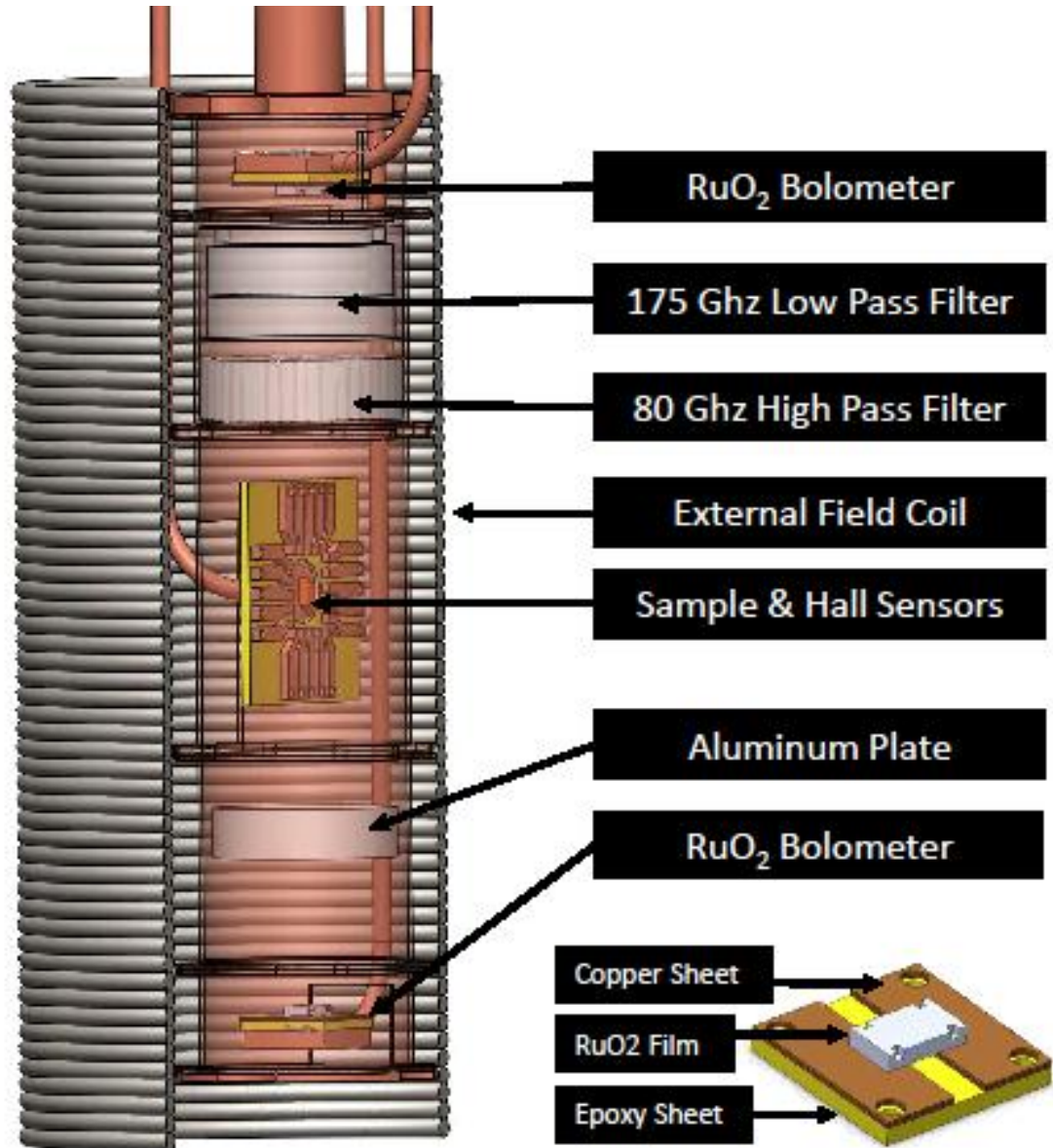


Questions:

1. Is there an EM radiation? What is its frequency?
2. How does the magnetization depend on location on the sample? Is there an Avalanche in Fe_8 ?

Experimental Setup.

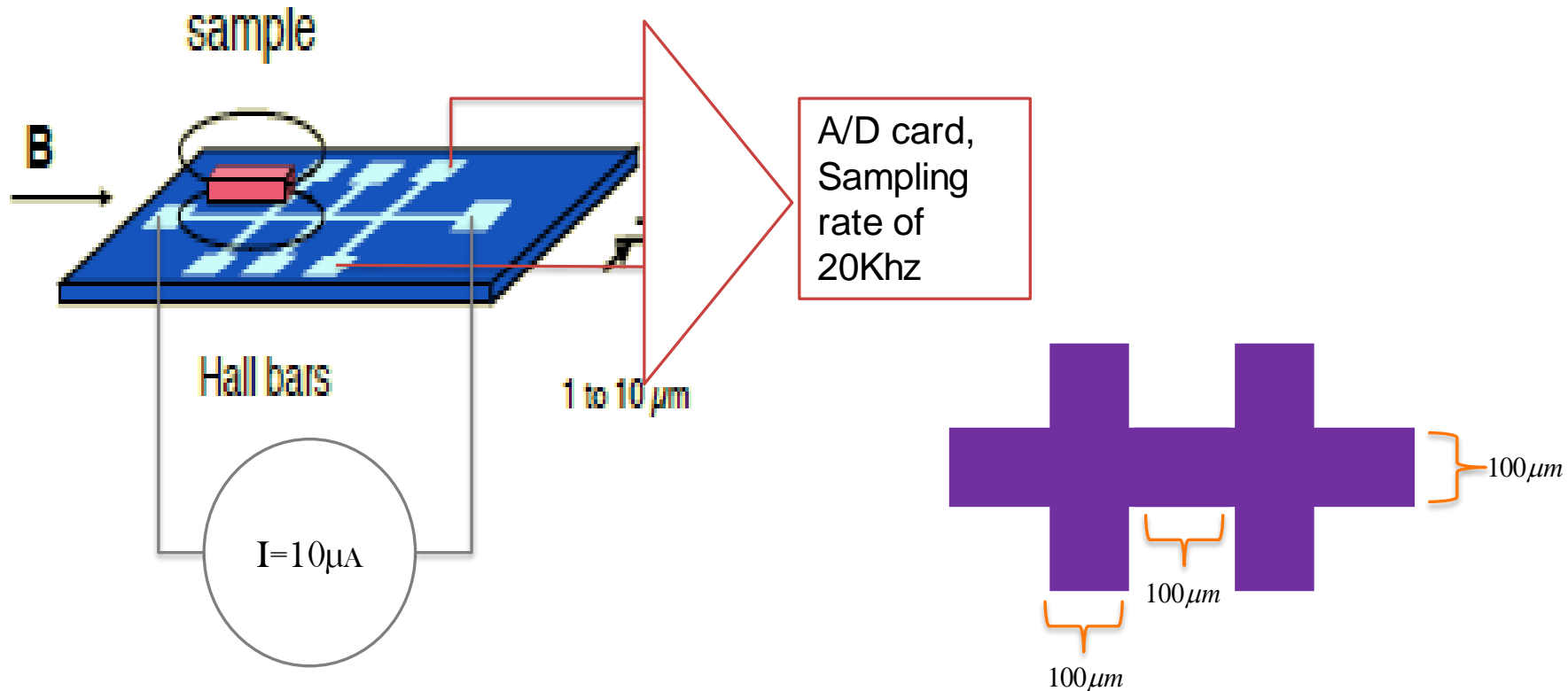
The Experiment conducted in Dilution Refrigerator at 100mK



Experimental Setup.

Hall Sensors - Form Eli Zeldov , Weitzman institute.

- Two-dimensional electron gas (2DEG) formed at the interface of GaAs/AlGaAs heterostructures.



Experimental Setup.

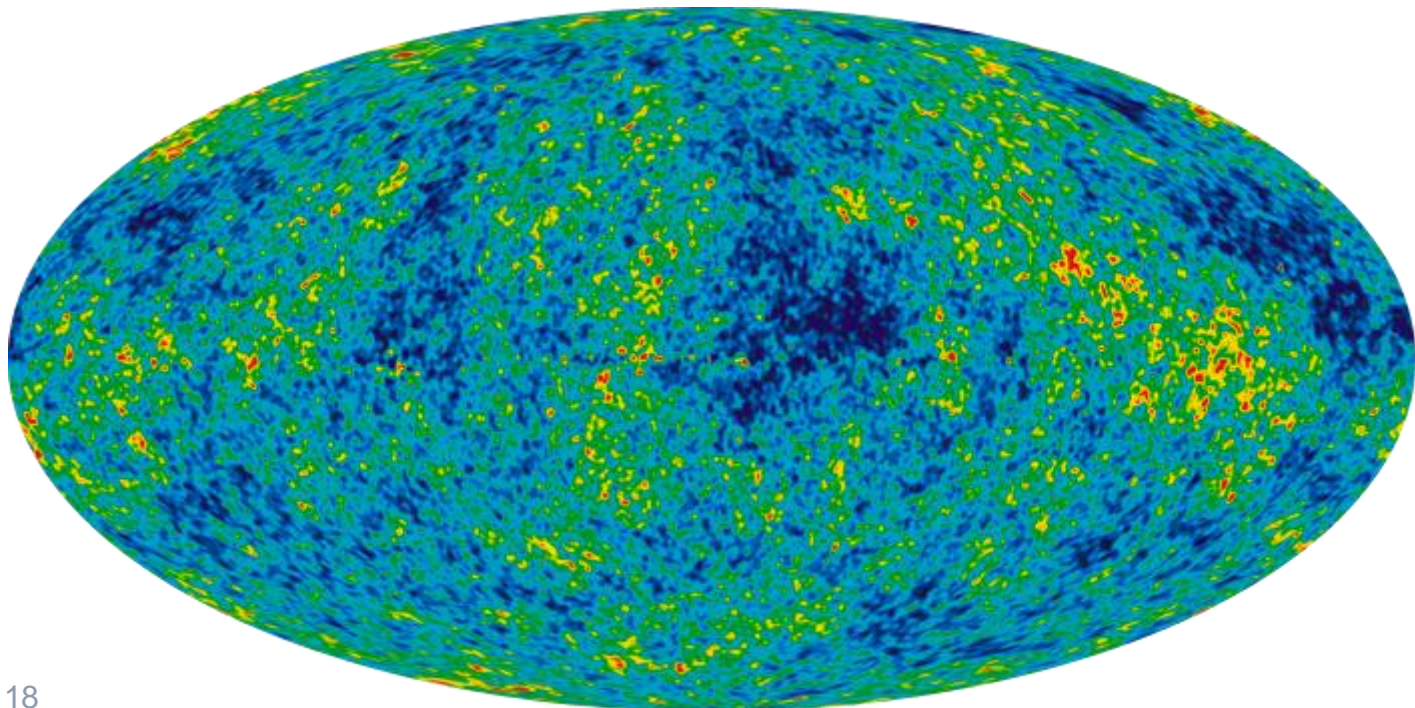
Filters

Used for cosmic background radiation measurement.

From Shaul Hanany – University of Minnesota

MAXIMA

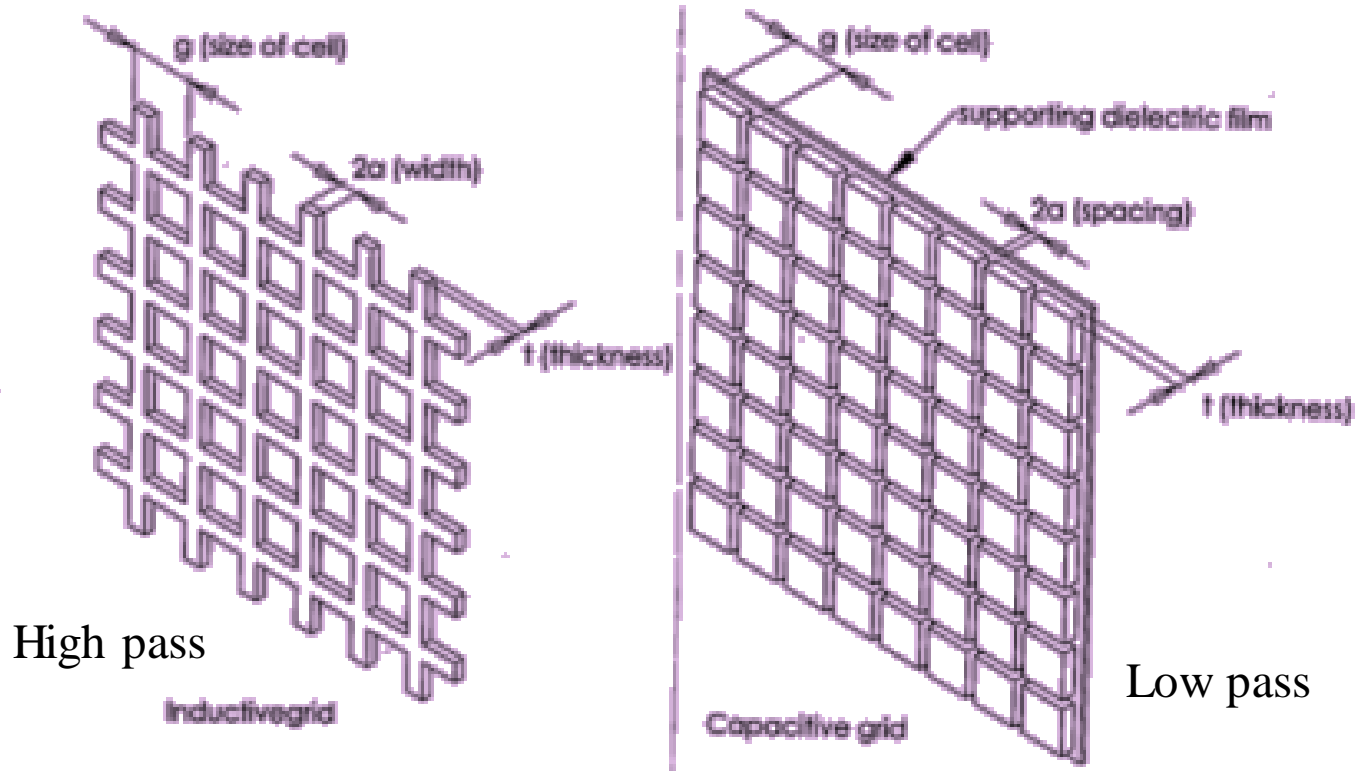
Millimeter Anisotropy eXperiment IMaging Array



Experimental Setup.

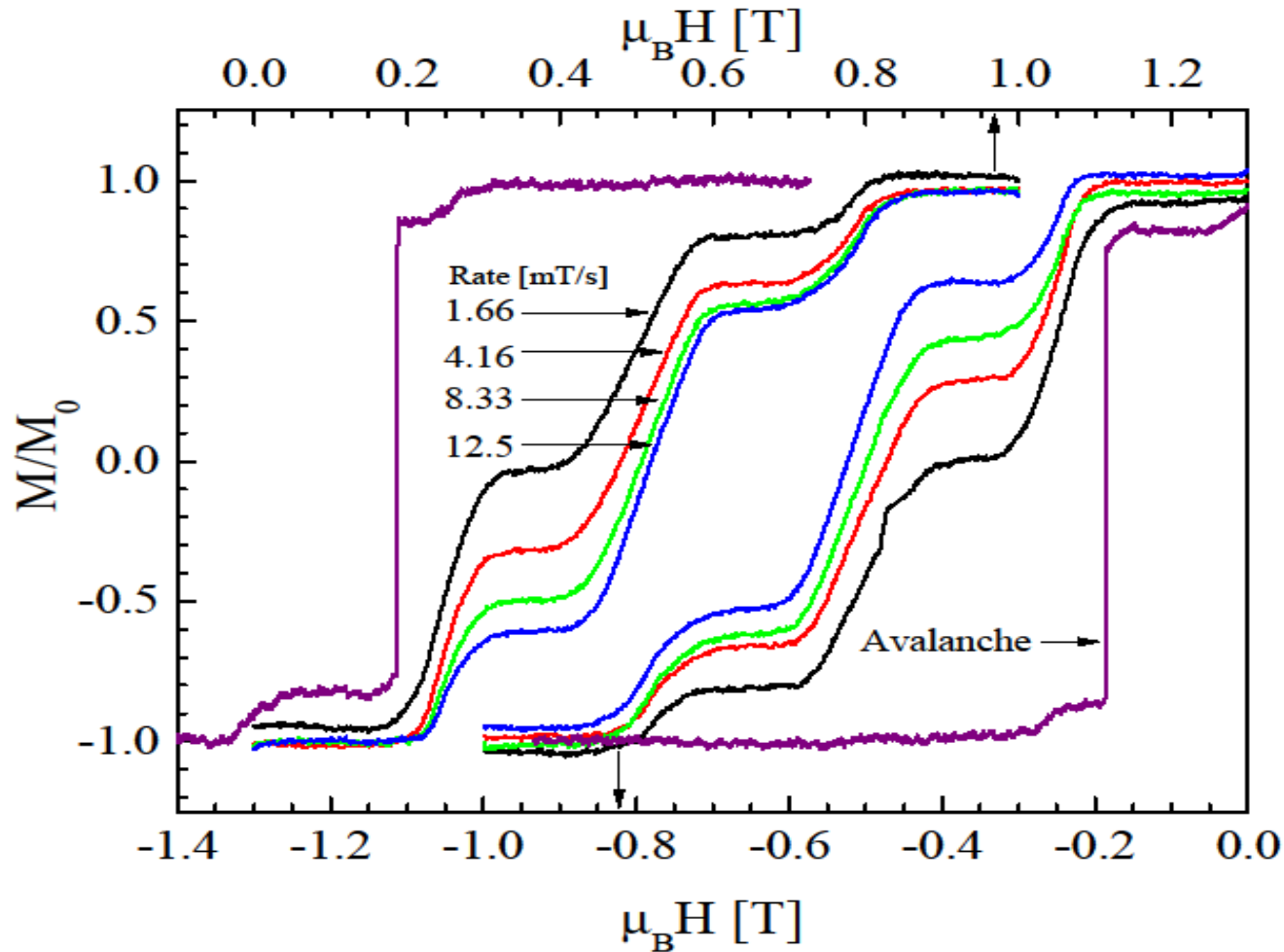
Filters

- In an inductive grid the metal is continuous, so at low frequency it must reflect all the incident wave (high pass).
- Because the grids are complements of each other, a capacitive mesh acts as a low pass filter.



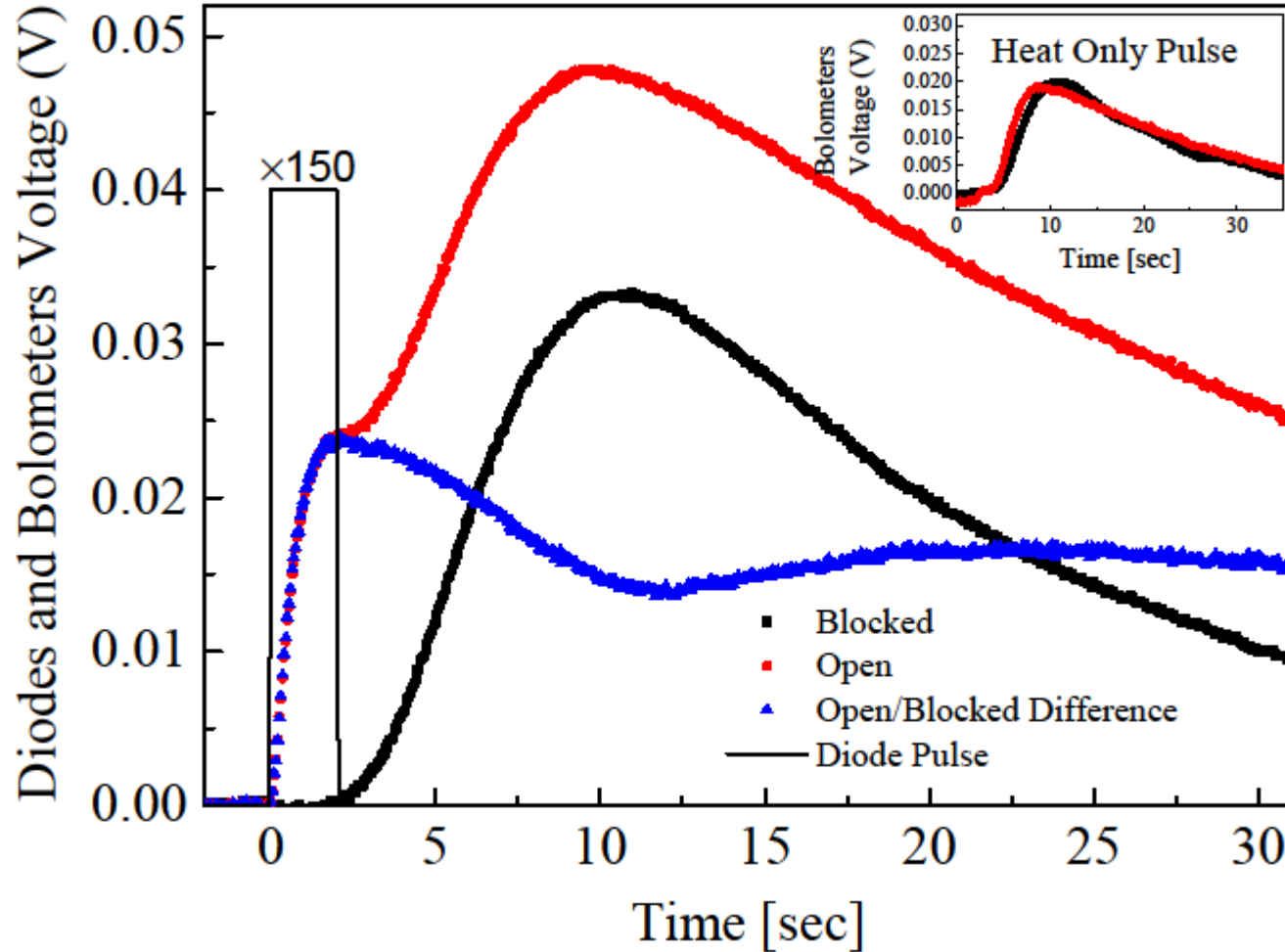
Results – Plan A

Magnetization Measurements



- Two types of relaxation are detected.
- Avalanche is new, Staircase is similar to previous work.
- Magnetic sweep rate dependence is seen.

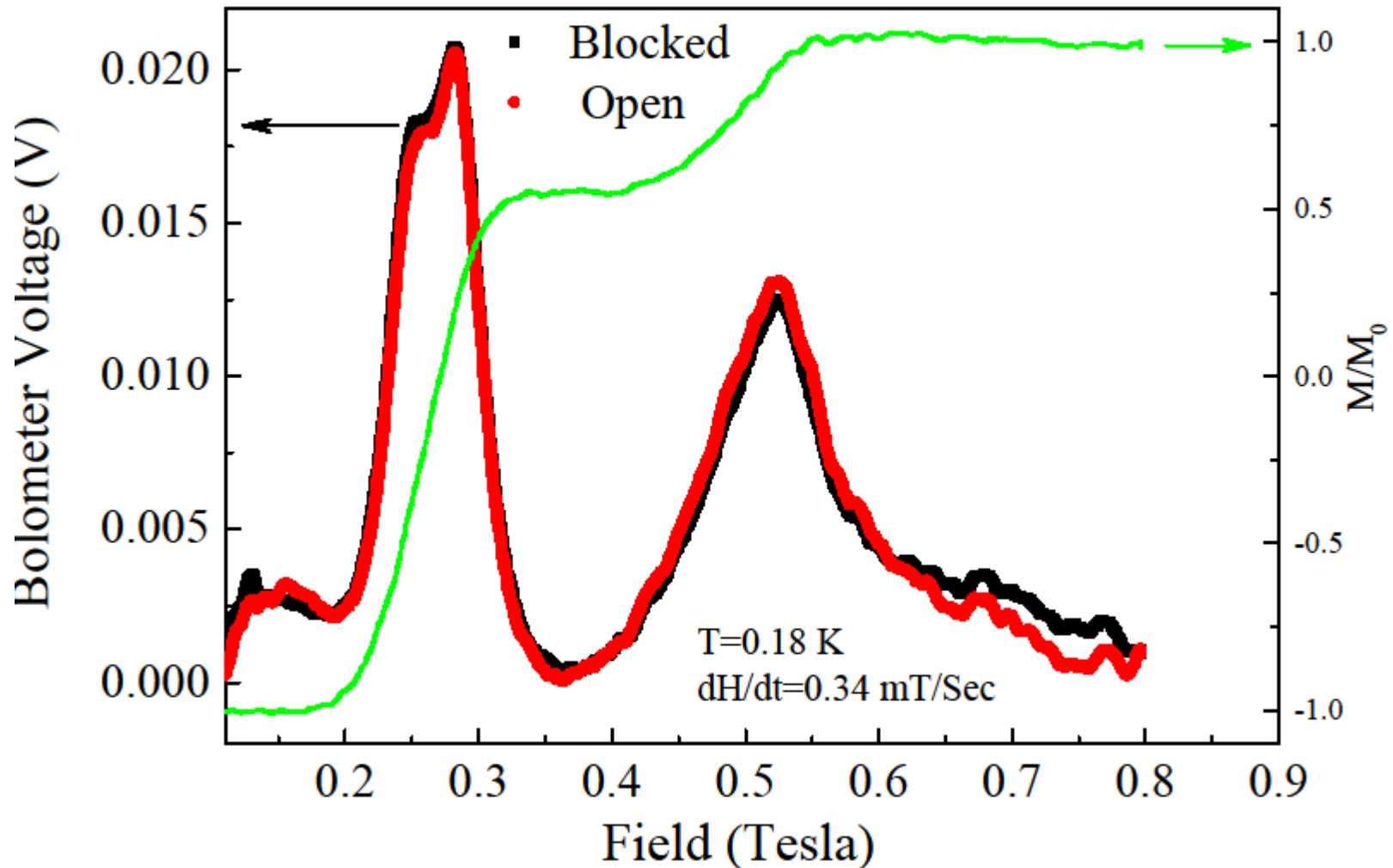
Radiation Test Experiment



Setup was tested by replacing the sample with 2 LEDs and a heater.

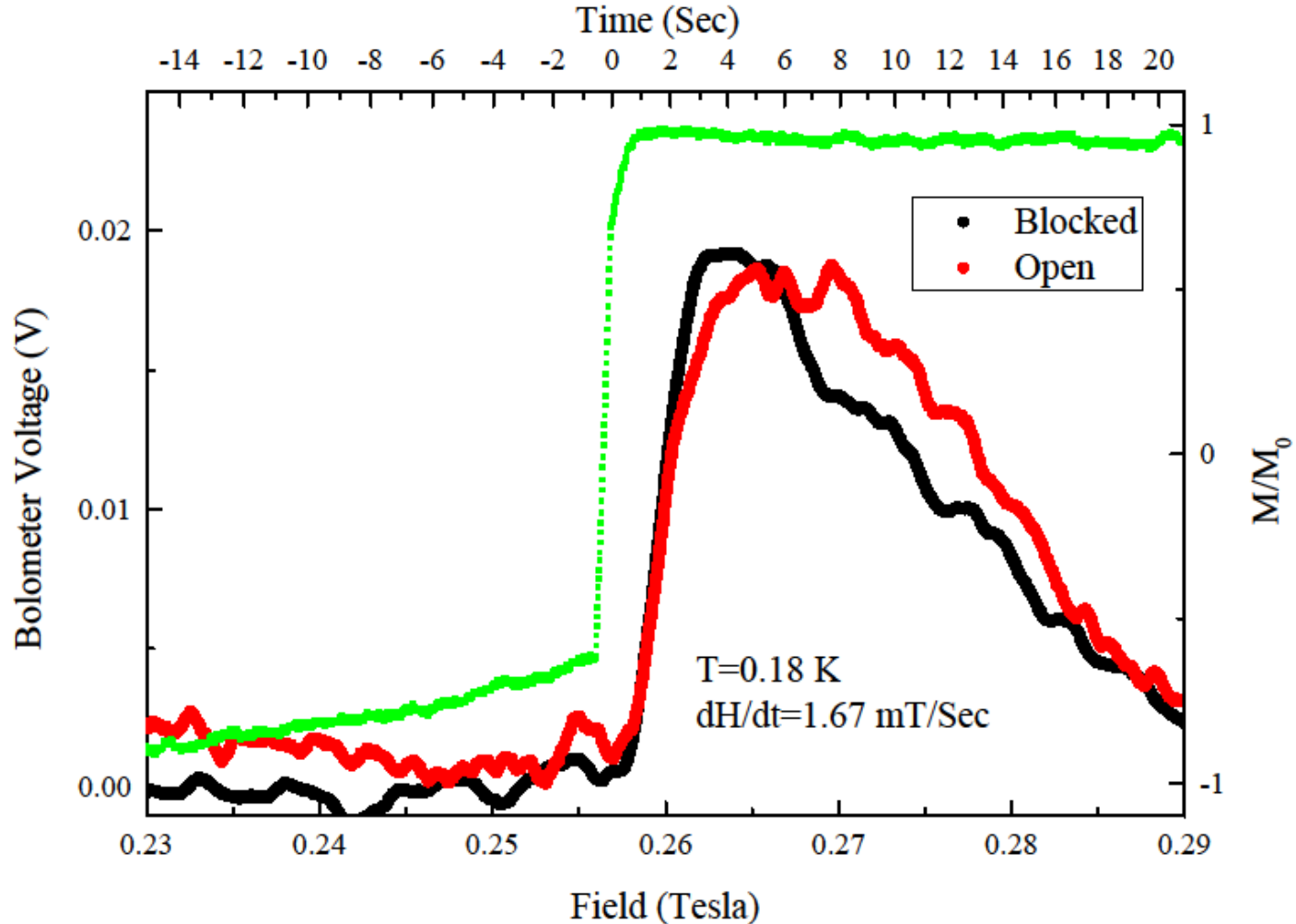
Bolometer which sees the Diode rises earlier and higher.

No Radiation has been Detected At the staircase regime ☹️



- Question: Why the second transition lasts longer than the first after tunneling have finished?

No Radiation has been Detected At avalanche regime ☹️



- Question: Where energy in avalanche? Same voltage change, much shorter time.

Plan A conclusions

- ✓ We found no evidence for release of electromagnetic radiation.
- ✓ Why the second transition lasts after the magnetization stops?
- ✓ Where is the energy in avalanche case?



Introduction to Magnetic Deflagration



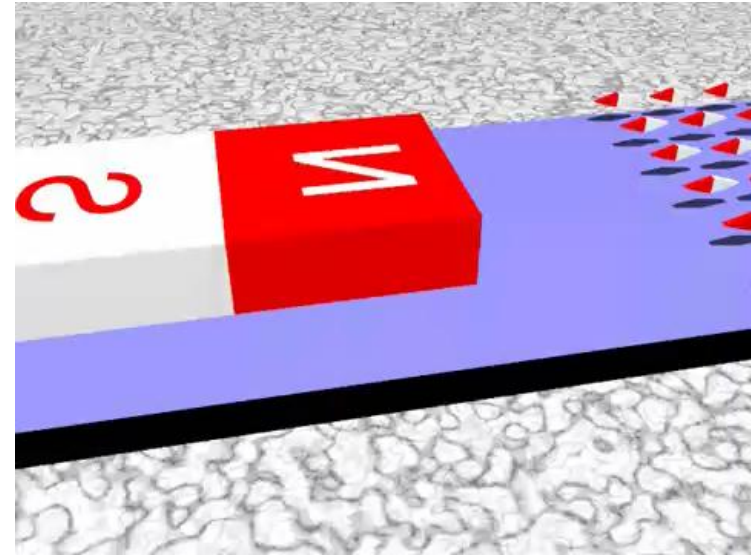
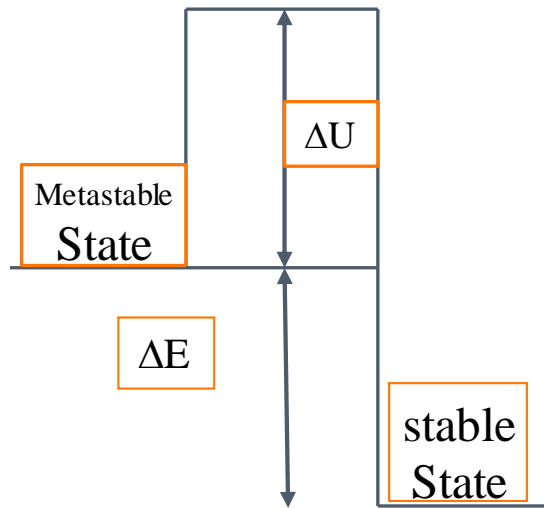
Magneto: The master of magnetism

What is a magnetic deflagration?

- Subsonic combustion propagating through heat transfer.
- Burning material heats the next layer of cold material and ignites it.
- In magnetic deflagration Zeeman energy playing the role of the chemical energy.
- Magnetic deflagration is nondestructive and reversible.



What is a magnetic deflagration?



➤ When a molecule makes a transition from the metastable state $S_z = \pm 9$ to the absolute energy minimum $S_z = \mp 10$, ΔE energy is released.

➤ This leads to local temperature change $\Delta T = \frac{\Delta E}{C_{ph}}$,

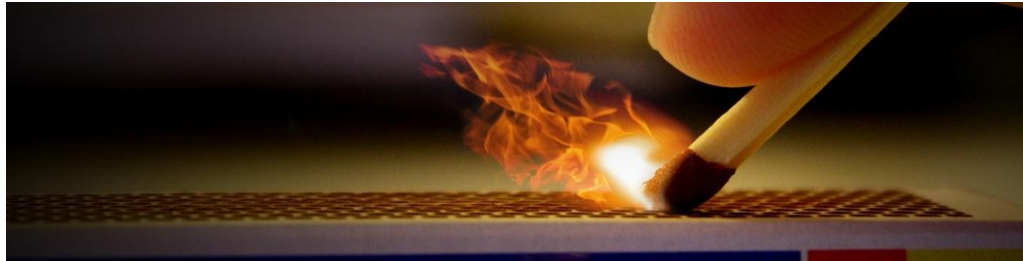
where is C_{ph} the phonon heat capacity per magnetic molecule.

What is a magnetic deflagration?

➤ Heat flows according to heat equation:

$$\frac{dT}{dt} = \nabla \kappa(T) \nabla T - \frac{\Delta E}{C_{ph}} \frac{dn}{dt}$$

$$\frac{dn}{dt} = -\Gamma(n - n^{eq})$$



➤ There are two possible scenarios:

- A. The heat defuses away – **No Ignition**. (Slowly moving match.)
- B. The heat cumulates, this rises the transition rates and causes **Deflagration**. (Match moves fast, heat creation is larger)

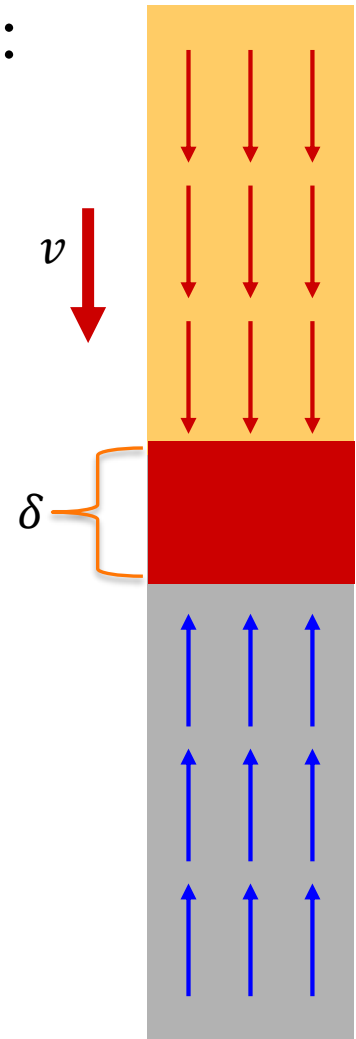
What is a magnetic deflagration?

When the transition rate of a single molecule $\Gamma(H_0, T_0)$
(Heat creation) exceeds a critical value:

$$\Gamma_c = \frac{8\kappa(T_0)k_B T_0^2}{U(H_0)\Delta E(H_0)n_i l^2}$$

Notice no sweep rate dependence

- $U(H_0)$ is the energy barrier.
- $\Delta E(H_0)$ is the energy released decaying to the ground state.
- n_i the initial population of the excited state.
- l characteristic length of the sample.
- T_0 the ignition temperature.



What is a magnetic deflagration?

There are two characteristic timescales:

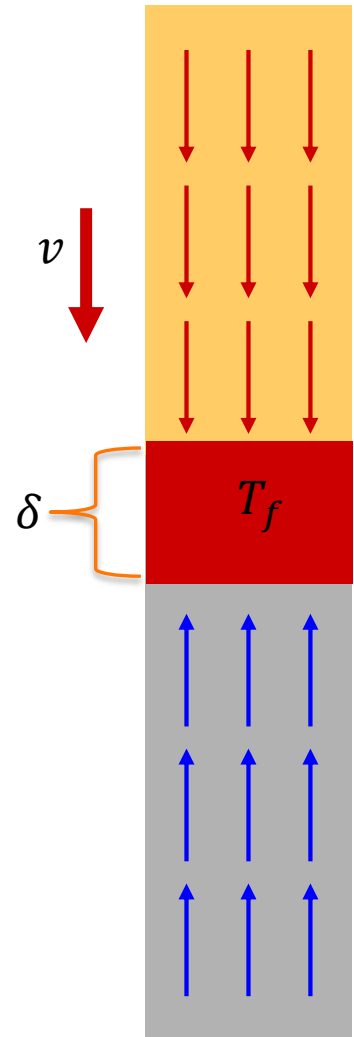
1. Thermal diffusion timescale: $\tau_d \simeq \frac{\delta^2}{\kappa}$
2. Burning timescale which is the chemical reaction timescale given by: $\tau_b = \tau_0 \exp\left(\frac{U(H)}{k_B T_f}\right)$

The width of the front δ , can be approximated by equalization the two timescales $\tau_d \simeq \tau_b$ leading to:

$$\delta = \sqrt{\kappa \tau_b}$$

The front propagates at a characteristic speed which equals to the width over the burning rate:

$$v \simeq \sqrt{\frac{\kappa}{\tau_b}} = \sqrt{\frac{\kappa}{\tau_0}} \exp\left(-\frac{U(H)}{2k_B T_f}\right)$$



Hall Signal Calculation

Can we calculate the deflagration front width from our data?

Magnetic dipole: $\vec{B} = \frac{\mu_0}{4\pi} \left(\frac{3\vec{r}(\vec{\mu} \cdot \vec{r})}{r^5} - \frac{\vec{\mu}}{r^3} \right)$

Assuming the sample is on XY plane with the magnetization pointing in \hat{x} direction: $\vec{\mu} = m dx dy dz \hat{x}$, m is dipole density.

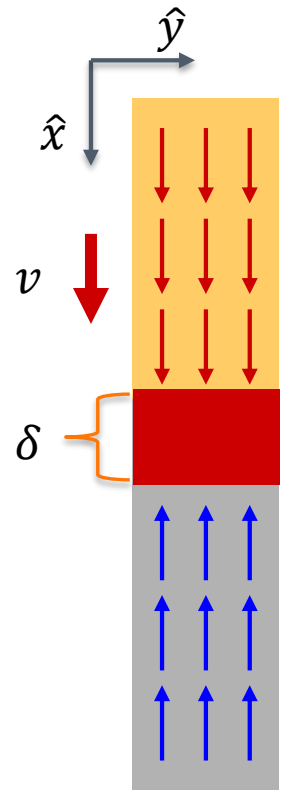
$$-L < x < x'(t)$$

$$-w < y < w$$

$$d_1 < z < d_2$$

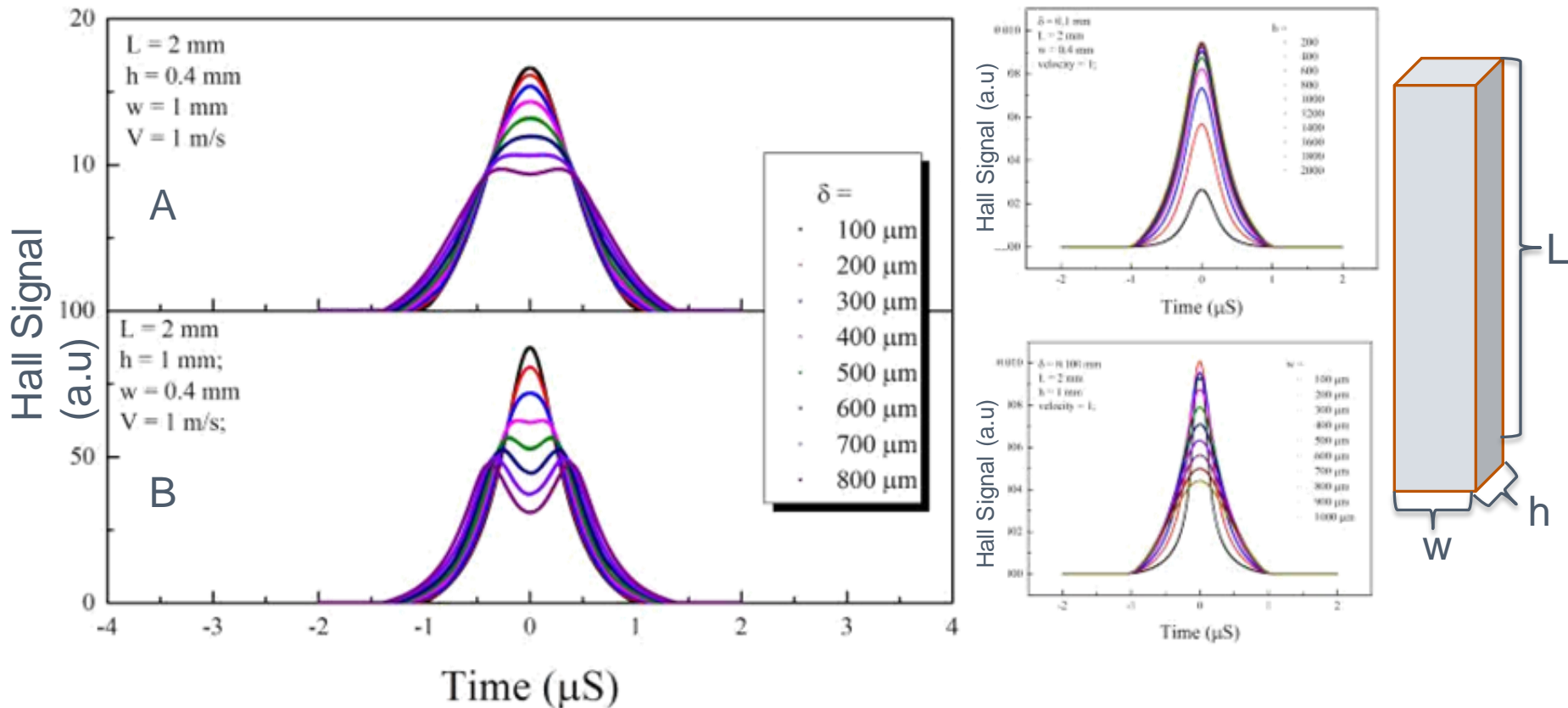
The Field in z direction can be calculated analytically:

$$B_z = \left[\text{Arctan} \left(\frac{\sqrt{w^2 + x^2 + d_2^2}}{w} \right) - \text{Arctan} \left(\frac{\sqrt{w^2 + x^2 + d_1^2}}{w} \right) \right]_{x=-L}^{x=x'(t)}$$



Hall Signal Calculation

Integration over the Hall Sensor for parameters similar to the experiment:



Things that can split the peak:

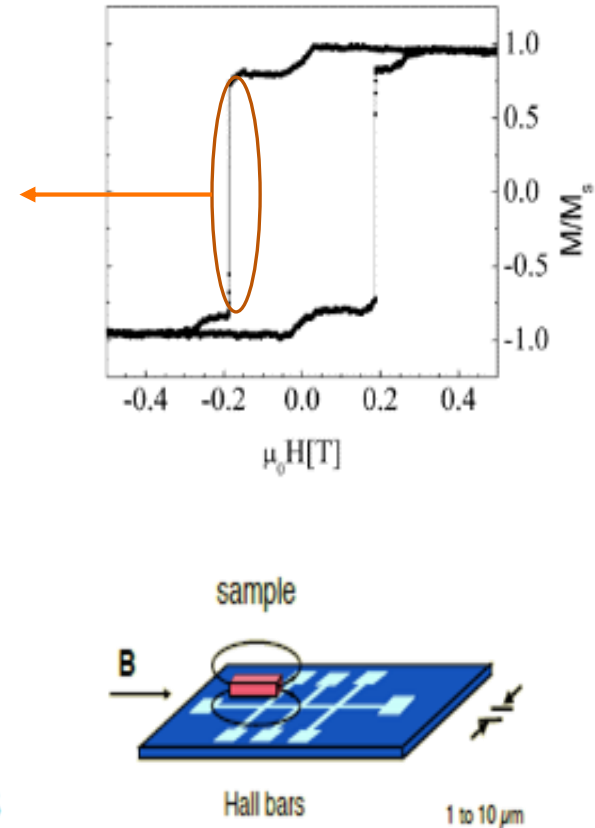
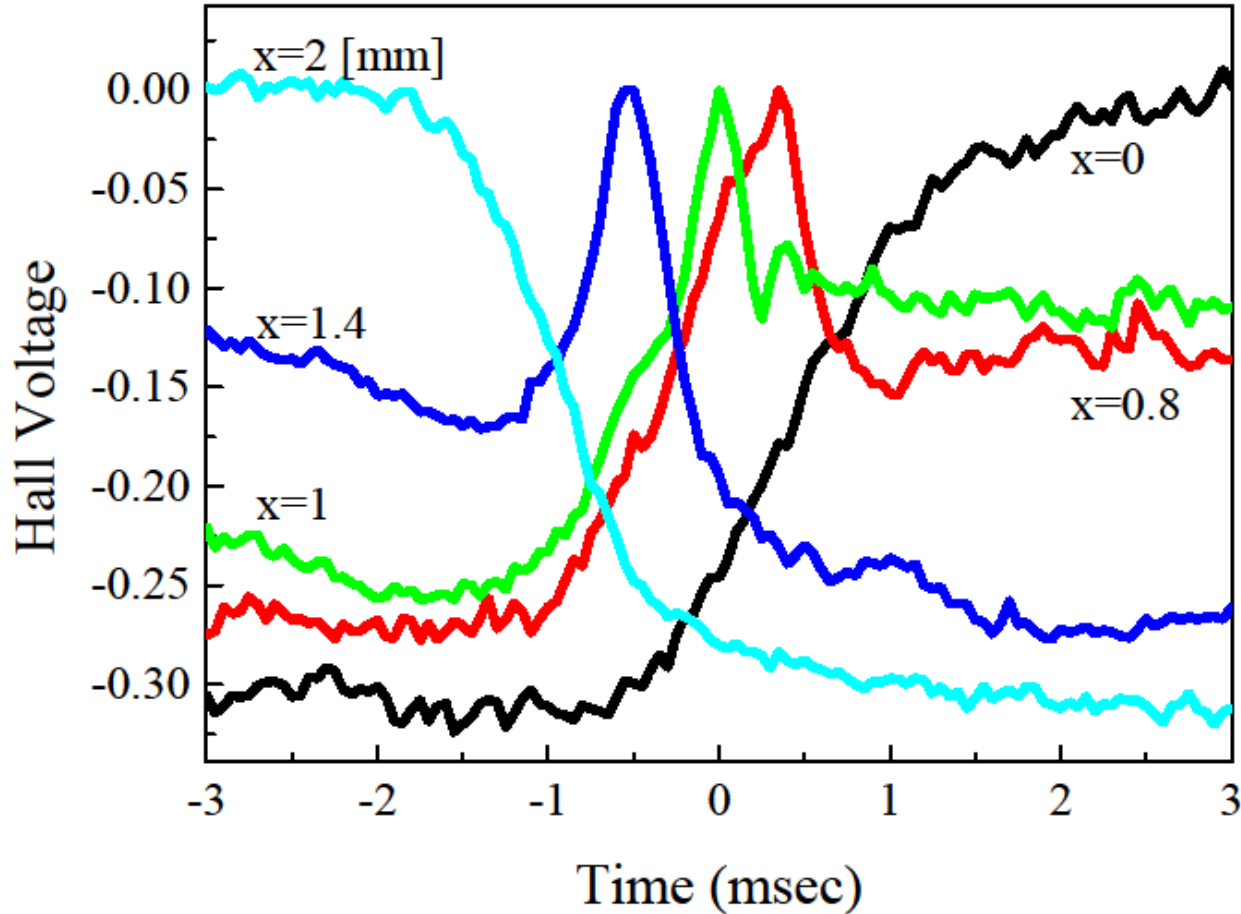
A. Deflagration front bigger than $\sim 700 \mu\text{m}$.

B. The deflagration runs through a strip smaller than the sample.

Results – Plan B

Results with Deflagration

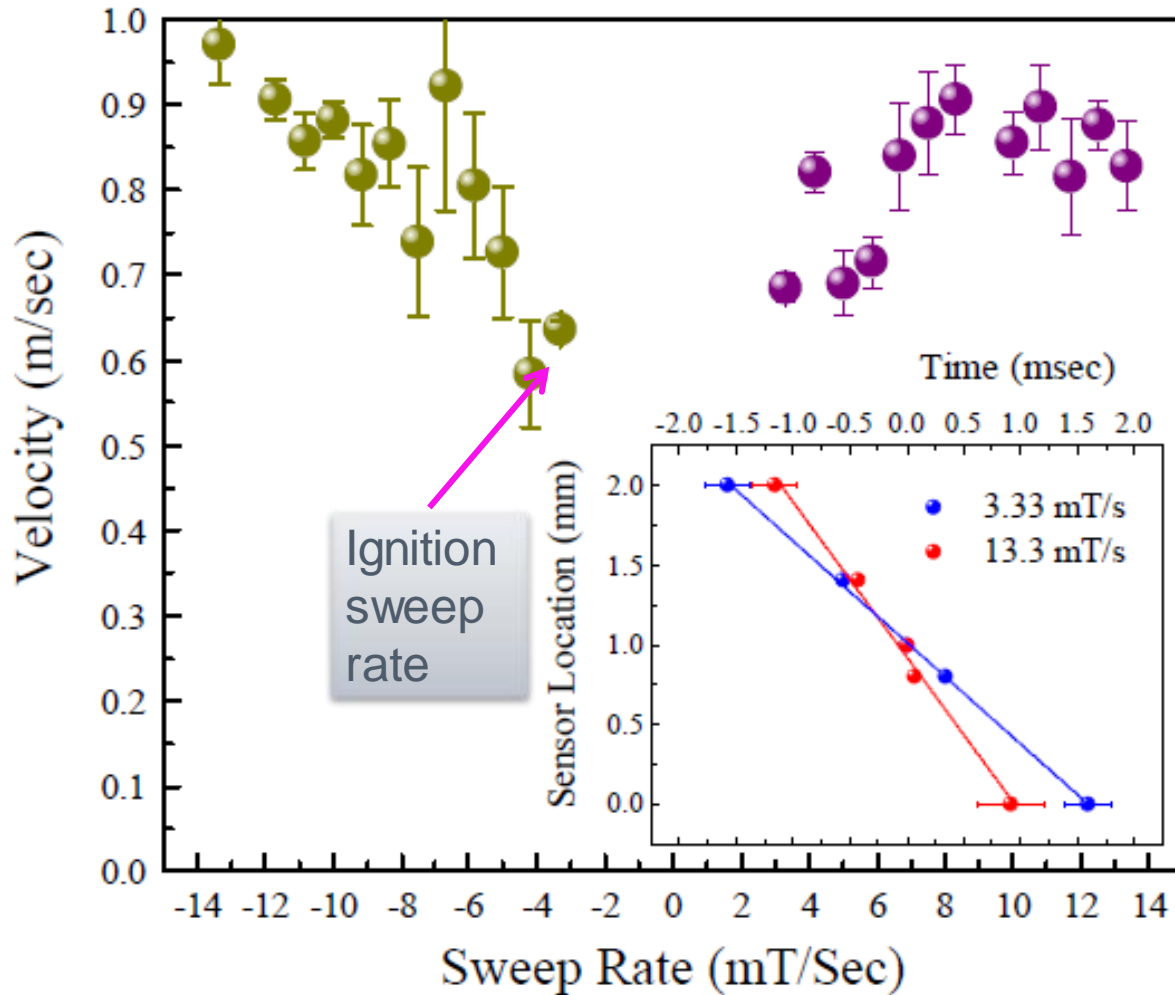
Above the critical sweep rate:



- The Deflagration ignites each time at first resonant field.
- The evolution of the peaks and cusps provide the avalanche propagation velocity.

Results

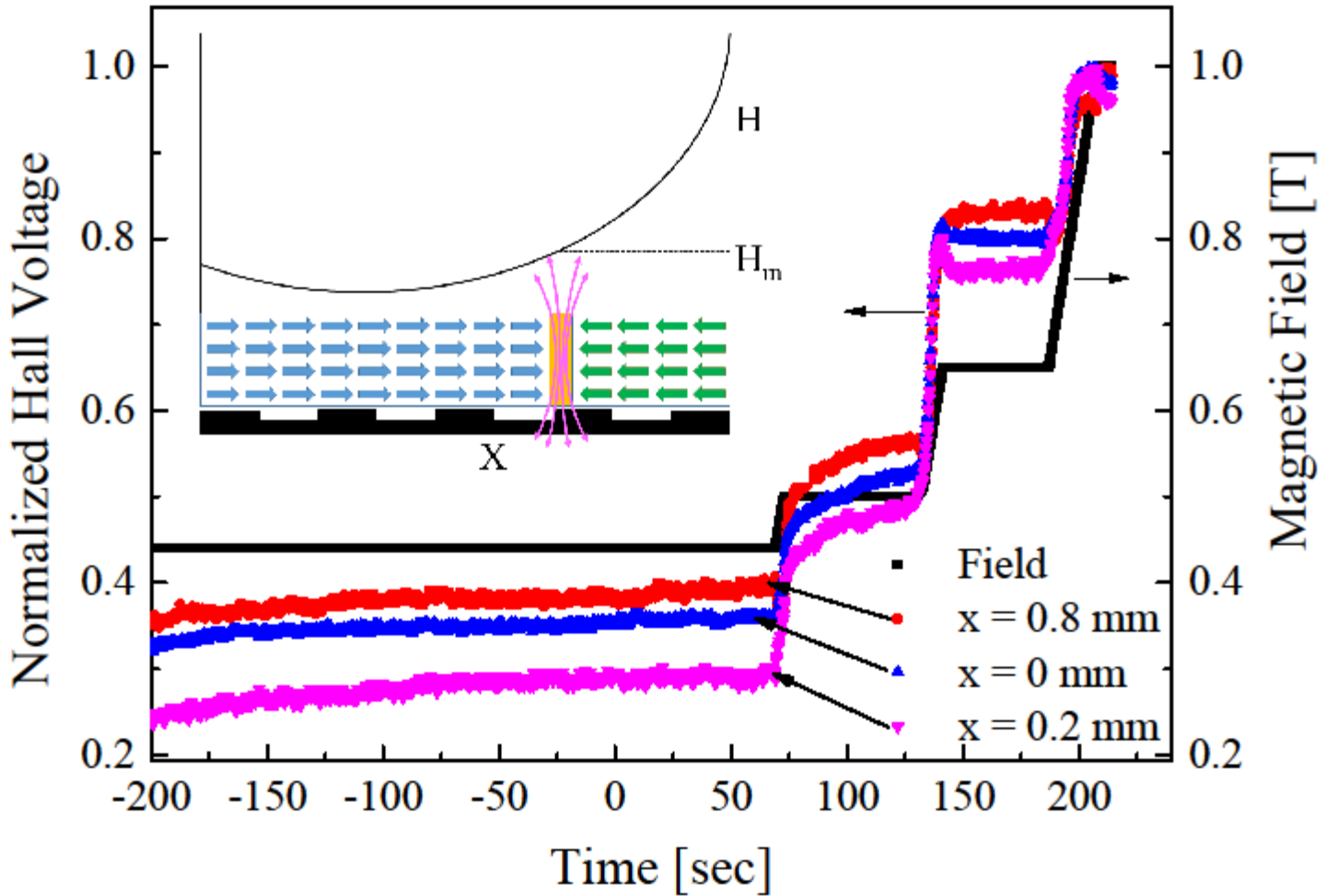
Deflagration velocity as a function of magnetic field sweep rate
No theoretical treatment



$$v = 0.5 - 1 \text{ m/sec}$$

Avalanche velocity as a function of magnetic field sweep rate at zero gradient. The field is swept from positive to negative and vice versa.

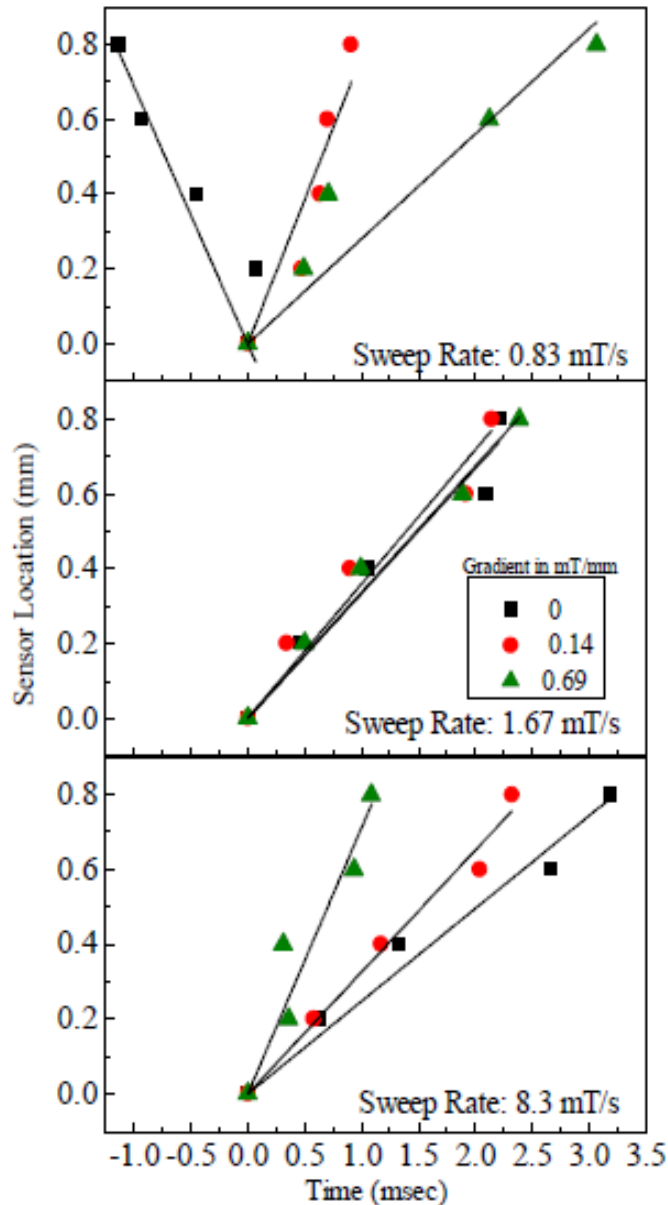
Result without Deflagration



- If the symmetry point, we might have a front of resonant area.
- In this case, pausing the field sweep will stop the magnetization evolution.
- For a typical sweep rate this speed is $V_m \sim 1.5 \times 10^{-4}$

Results

Different Gradients – No theoretical treatment



- The velocity not caused by inhomogeneity of the field.
- The velocity changes its size and direction as a function field gradient.
- Field Gradient can change the ignition point, the critical values and the propagation velocity, by changing the amount of molecules at resonance.

Deflagration Velocity Calculation

The velocity is given by:

$$v = \sqrt{\frac{\kappa}{\tau_0}} \exp\left(-\frac{U(H)}{2k_B T_f}\right)$$

Known:

$$\tau_0 = 3.4 \cdot 10^{-8} \text{ sec}$$

$$U(H) = 24.5 \text{ Kelvin}$$

$$v = 0.5 - 1 \text{ m/sec}$$

Not Know (yet):

$$\kappa = ?$$

$$T_f = ?$$

Thermal Diffusivity Measurements

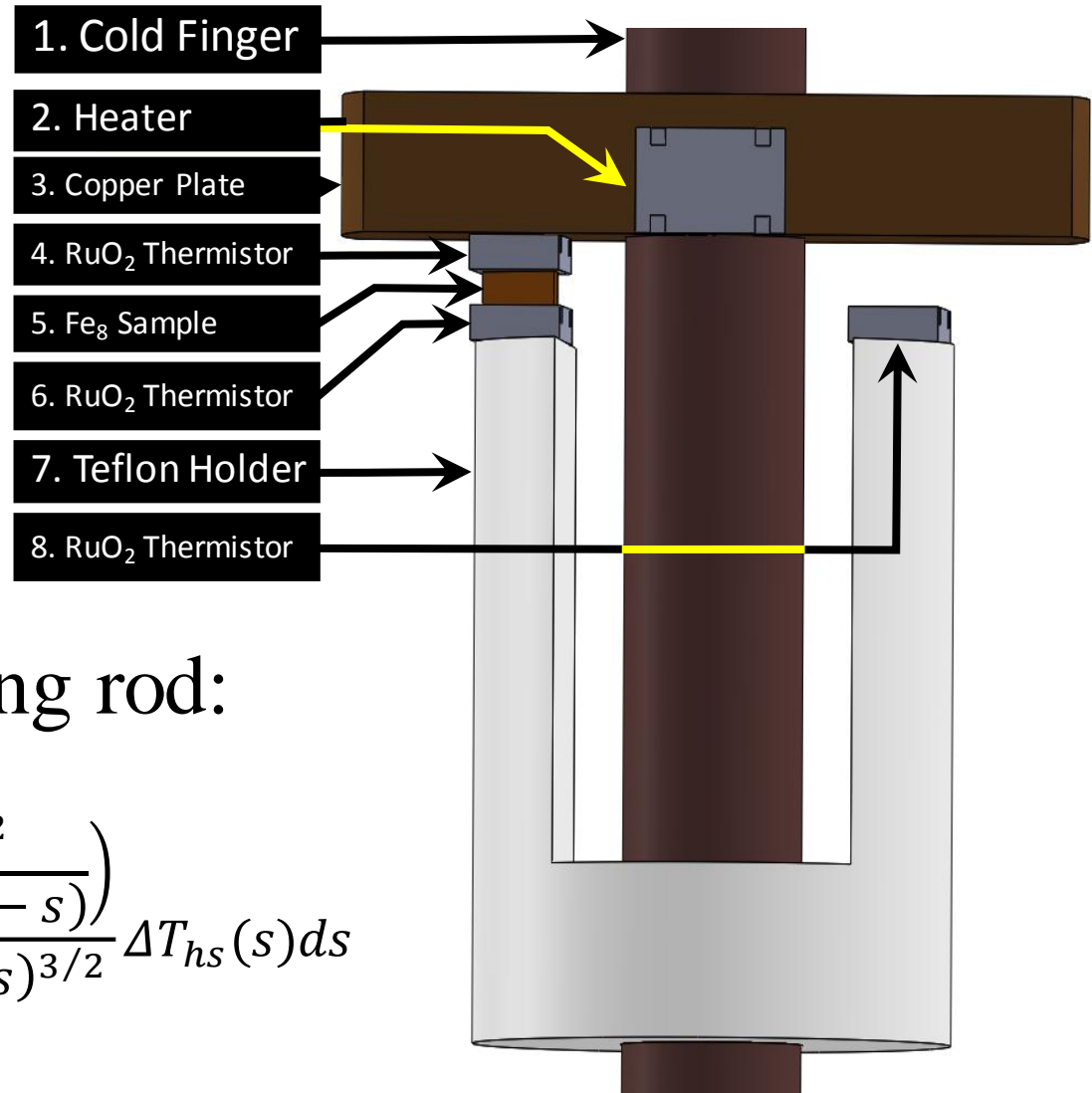
Experimental Setup

Thermal diffusivity is defined via the heat equation:

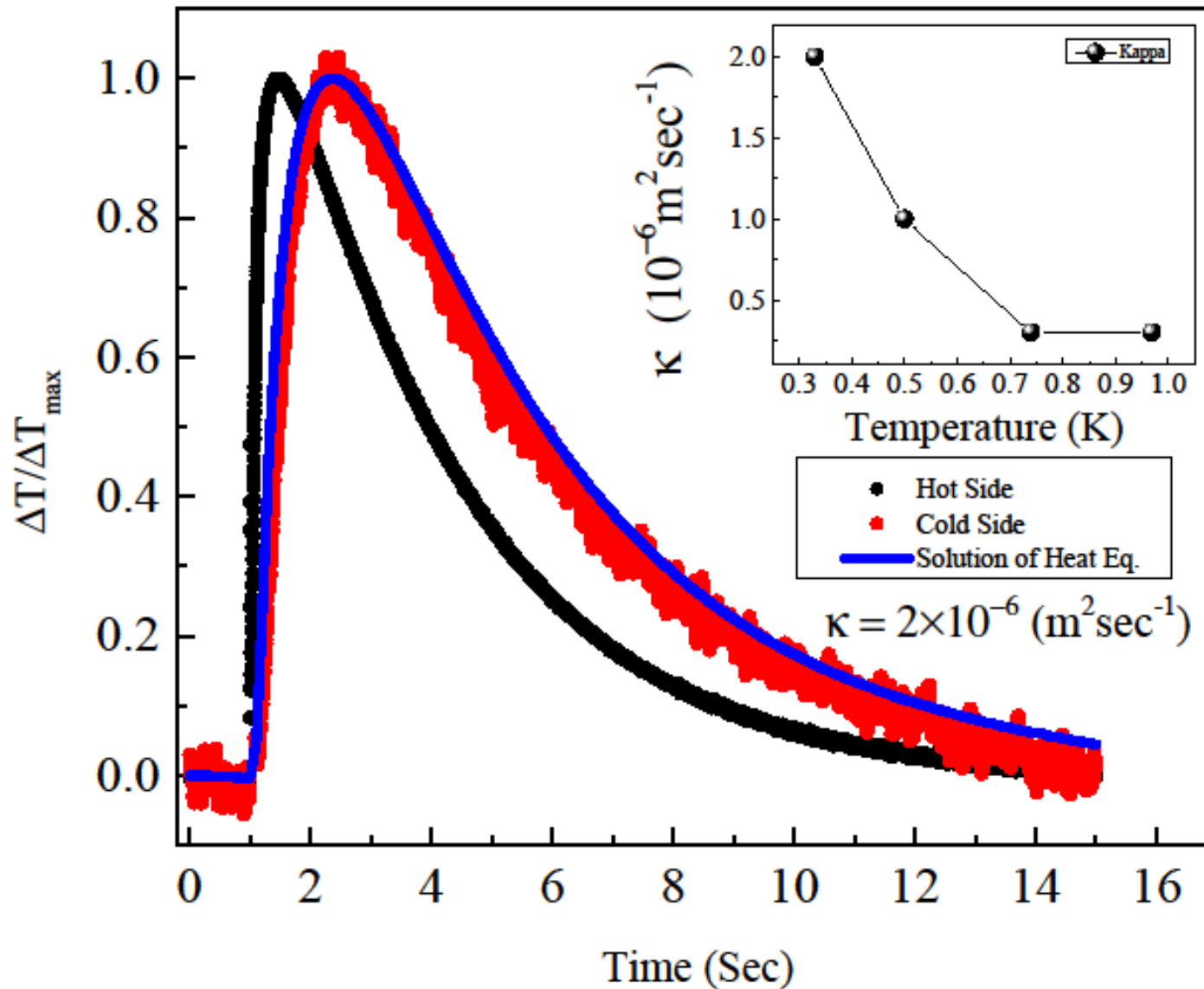
$$\frac{\partial T}{\partial t} = k \nabla^2 T$$

Solution for 1D long rod:

$$\Delta T_{cs}(t) = c \int_0^t \frac{x \exp\left(-\frac{x^2}{4\kappa(t-s)}\right)}{(4\pi\kappa)^{1/2}(t-s)^{3/2}} \Delta T_{hs}(s) ds$$



Thermal diffusivity measurements



Calculation of Flame Temperature

$$T_f = \frac{U}{k_B \ln\left(\frac{\kappa}{v^2 \tau_0}\right)} = \frac{24.5}{\ln\left(\frac{2 \cdot 10^{-6}}{0.6^2 \cdot 3.4 \cdot 10^{-8}}\right)} = 4.8K$$

- ✓ Seems a bit high for a sample at 180-300 mK.
- ✓ But still is lower than the barrier.

Measuring T_f will be Maayan's Project

Results Summary

- ✓ We found no evidence for release of electromagnetic radiation.
- ✓ We found deflagration in Fe₈ crystals.
- ✓ The deflagration front velocity is of the order of 1 m/sec and is sensitive to field gradients and sweep rates.
- ✓ We also measured the thermal diffusivity of Fe₈ and predicted The flame temperature.

Open Questions

- ❖ The heat released during avalanche is smaller than during staircase magnetization change, while it should be otherwise. Where did this energy go?
- ❖ Why is there critical sweep rate for deflagration?
- ❖ Why deflagration Velocity increases with sweep-rate?
- ❖ What is the field gradient dependence?
- ❖ Estimated flame temperature is around 5K which is pretty high and still lower than the barrier.



I would like to thank to



Amit
Keren



Shaul
Hanany



Eli
Zeldov



Emil
Polturak



Misha
Reznikoiv



Gil Drachuck and Tal Kirzhner



Asher Space Research Institute for funding

Thank You
For Listening

Sensitivity

Time (Sec)

-14 -12 -10 -8 -6 -4 -2 0 2 4 6 8 10 12 14 16 18 20

