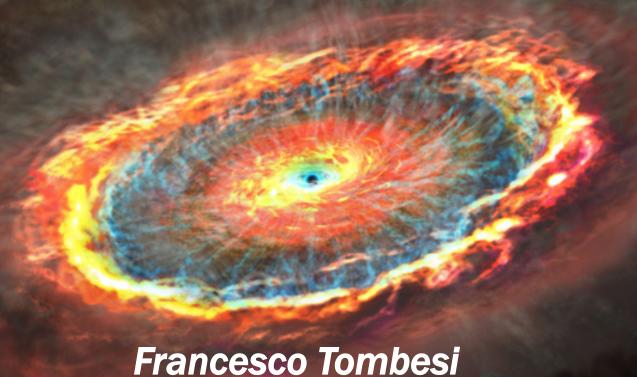
X-ray observations of UFOs and their implications for AGN feedback



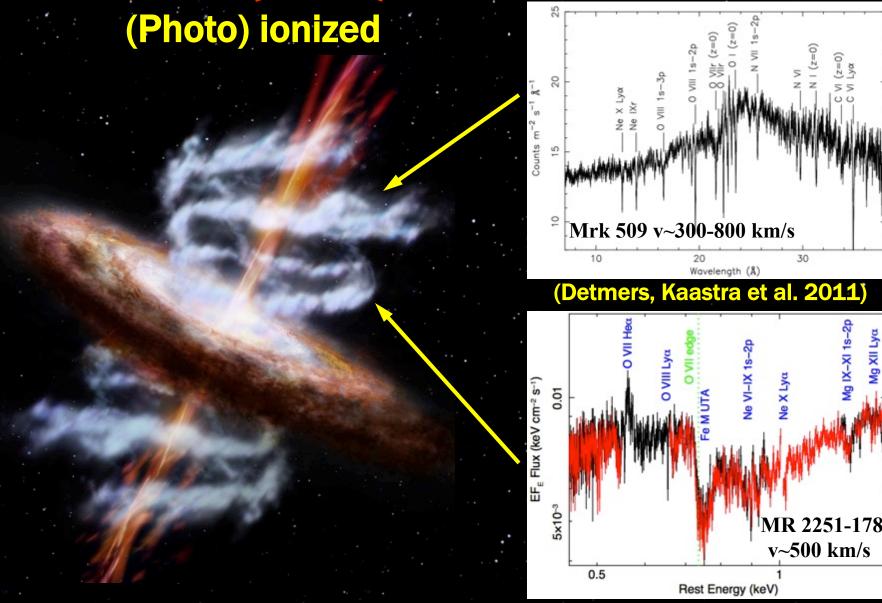
University of Rome, Tor Vergata

NASA - Goddard Space Flight Center

University of Maryland, College Park

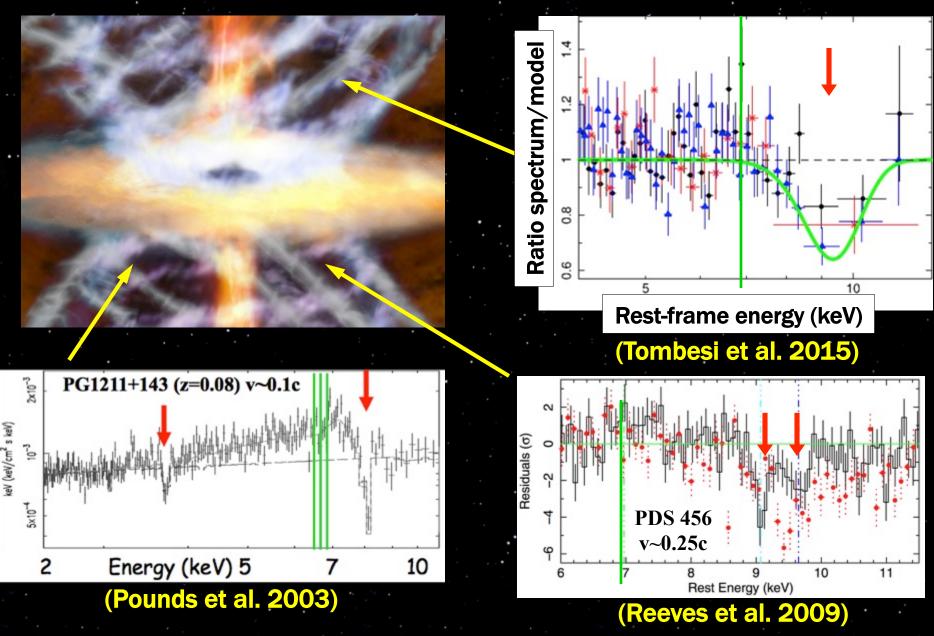
(credit JAXA's press release)

Warm Absorbers (WAs)



(Reeves et al. 2013)

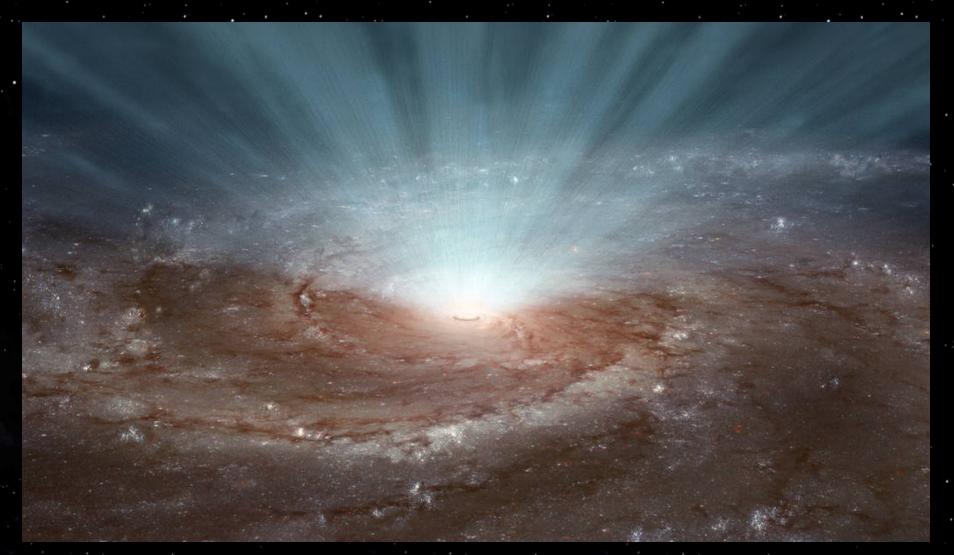
UltraFast Outflows ("UFOs")



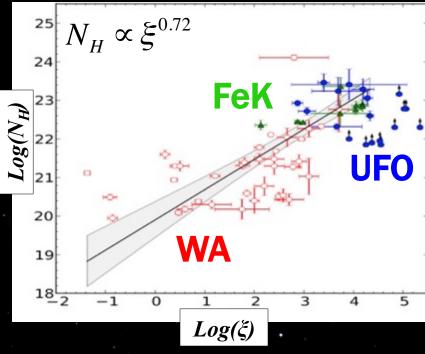
Some open questions on AGN winds:

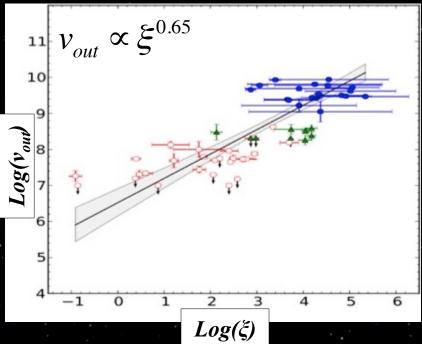
- 1. What is the relation between UFOs and warm absorbers?
- 2. What is the relation between UFOs and UV outflows?
- 3. What is the relation between UFOs and AGN jet?
- 4. Are UFOs energetic enough to drive AGN feedback?
- 5. Are there evidences for AGN wind feedback in action?

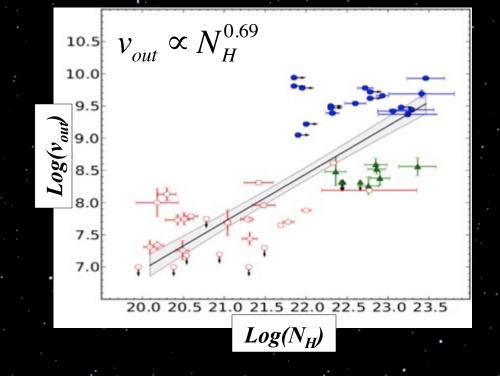
1 - What is the relation between UFOs and warm absorbers?



F. Tombesi - AGN Driven Winds, 22-26 May 2017, Technion, Israel



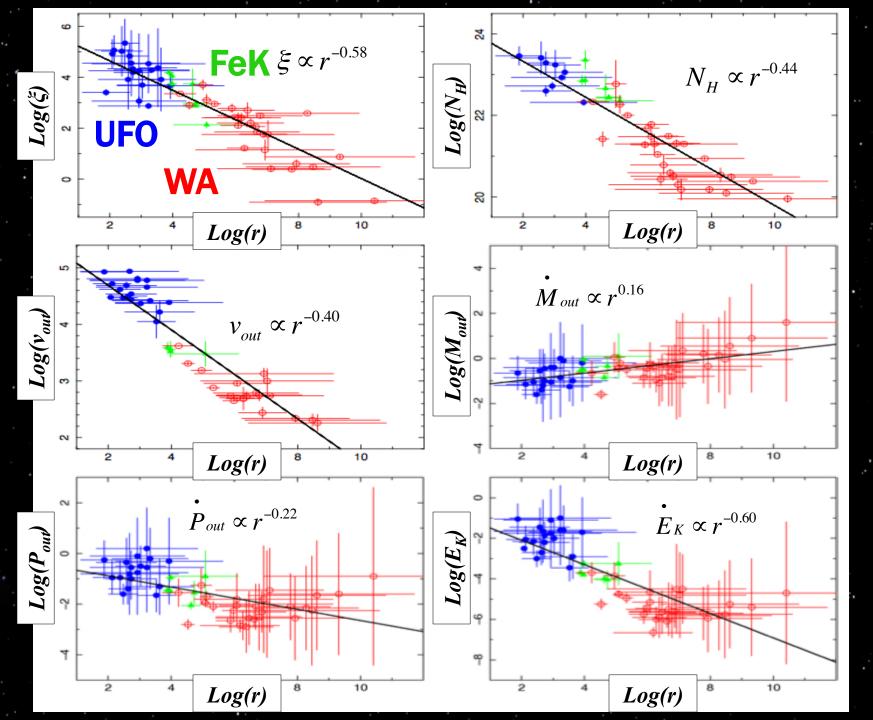




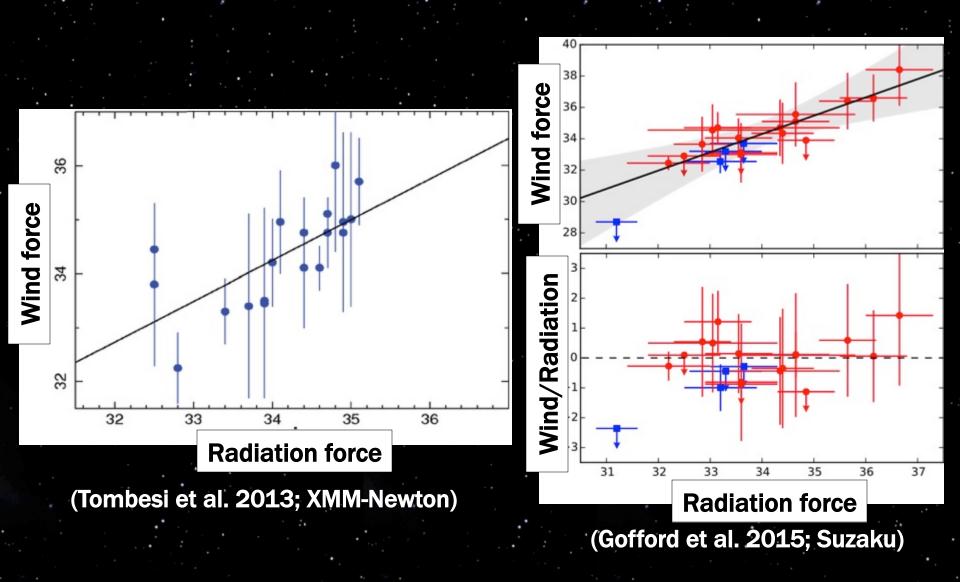
Comparison of UFOs and WAs in sample of 35 Seyfert 1 galaxies

- WAs found in >60% sources
- UFOs in >40%, >70% also WAs
- Profiles: n~1/r^{1.4}, v_{out}~1/r^{0.5}

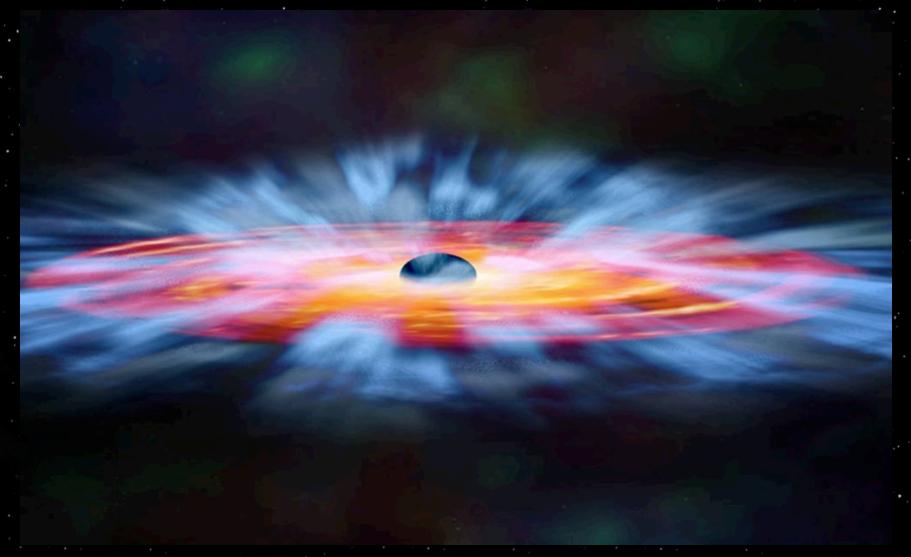
(Tombesi et al. 2013)



Correlation between UFO and AGN momentum rates

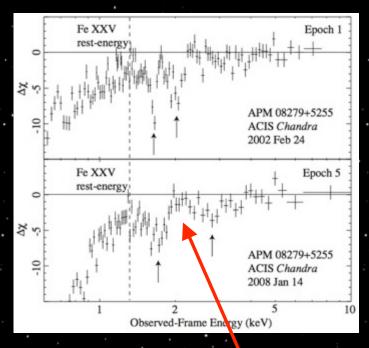


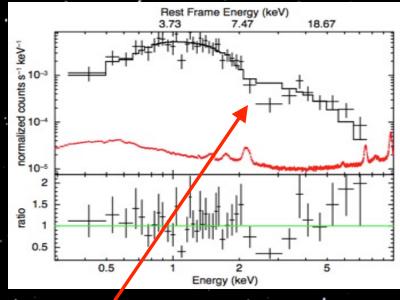
2 - What is the relation between UFOs and UV winds?



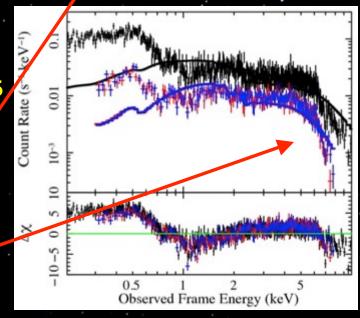
F. Tombesi - AGN Driven Winds, 22-26 May 2017, Technion, Israel

Ultrafast outflows in (high-z) BAL quasars

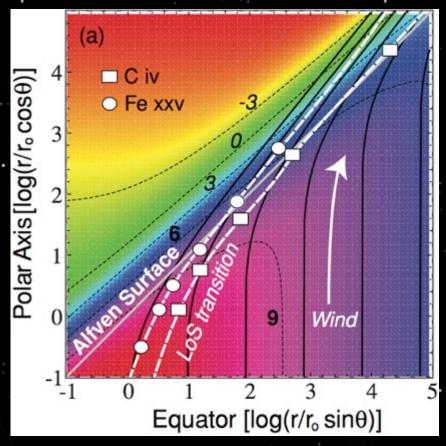




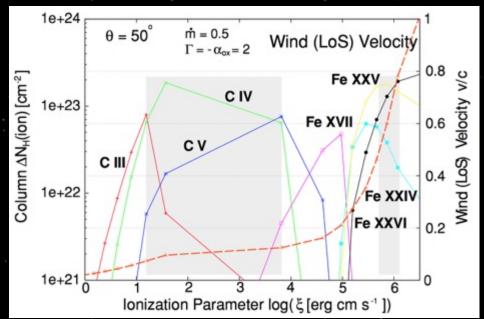
- Gravitationally lensed BAL QSO APM 08279 +5255 (z = 3.9), $v_{out} \sim 0.2$ -0.7c (Chartas et al. 2009)
- NAL QSO HS 1700+6416 (z= 2.7), v_{out}~0.1-0.6c
 (Lanzuisi et al. 2012)
- Mini-BAL QSO PG1126-041, v_{out}~16,500 km/s (Giustini et al. 2011)

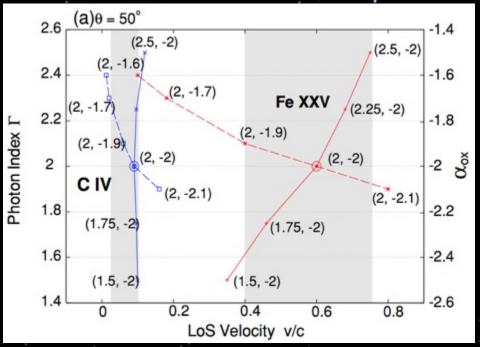


What is the relation between UFOs and BALs?

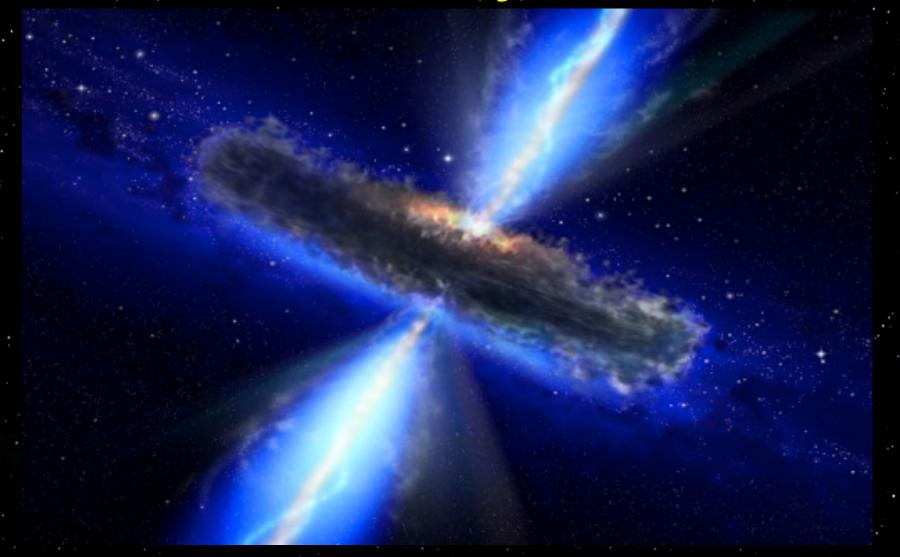


- Gravitationally lensed BAL quasar APM 08279+5255 (z = 3.9) C IV v~0.1c, Fe XXV v~0.5c
- MHD disk wind can explain both lines for α_{0x} ~-2 (Fukumura et al. 2010)



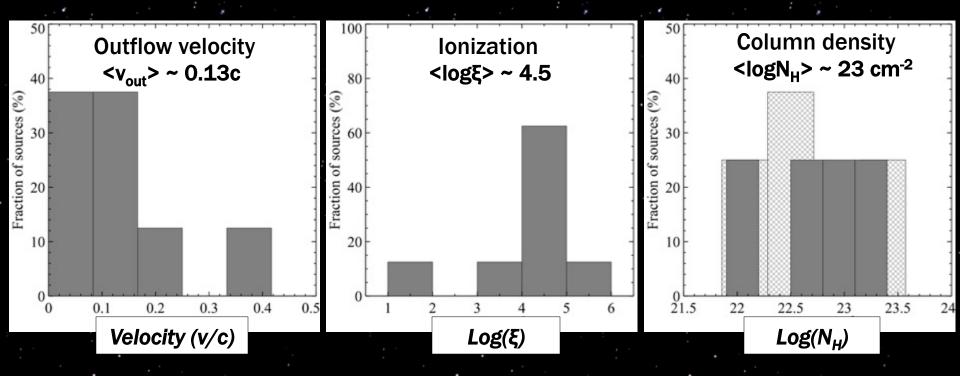


3 - What is the relation between UFOs and AGN jets?



F. Tombesi – AGN Driven Winds, 22-26 May 2017, Technion, Israel

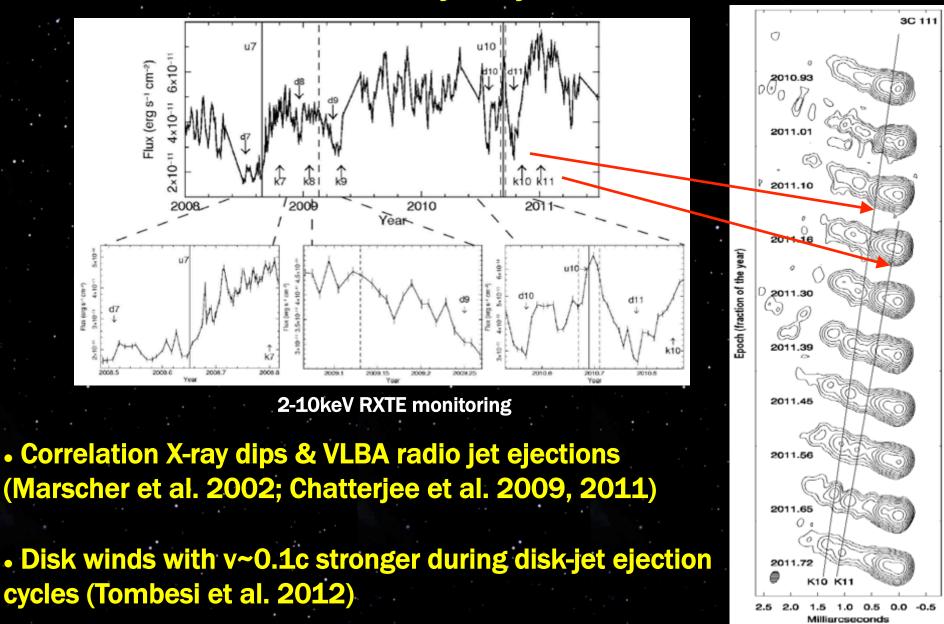
Ultra-fast outflows in radio-loud AGNs



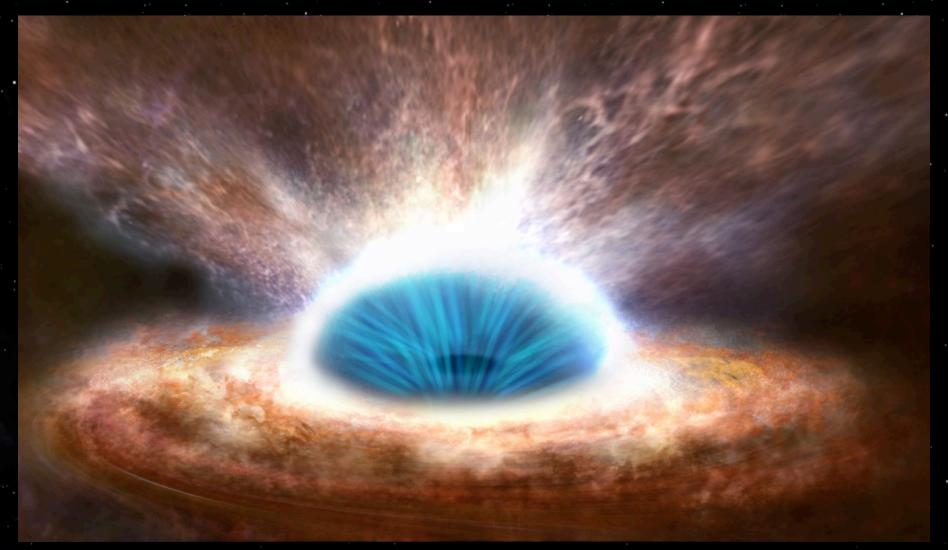
- Combining results with literature, UFOs in 7/26 (~30%) sources
- But only \sim 56% spectra have enough S/N, frequency of UFOs is $f=(50\pm20)\%$
- Similar to RQ AGNs: RQ/RL dichotomy does not apply to disk winds?

(Tombesi et al. 2014)

Disk, wind, and jet cycles in 3C 111

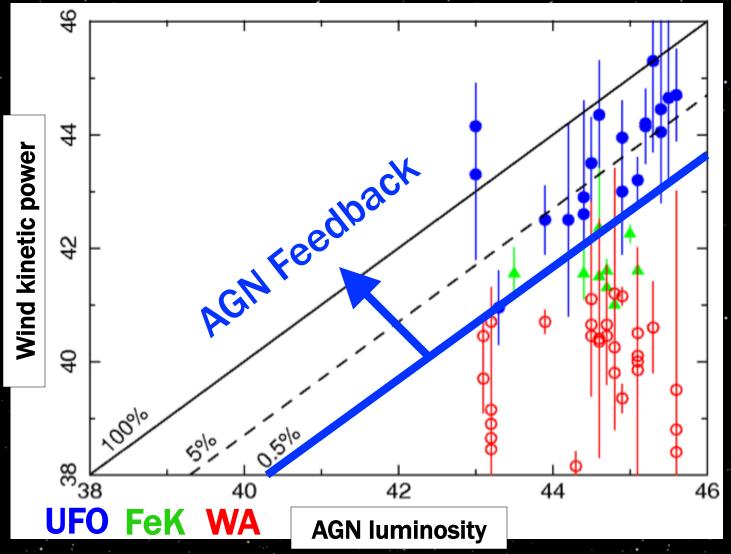


4 - Are UFOs energetic enough to drive AGN feedback?



F. Tombesi – AGN Driven Winds, 22-26 May 2017, Technion, Israel

How powerful are ultrafast outflows?

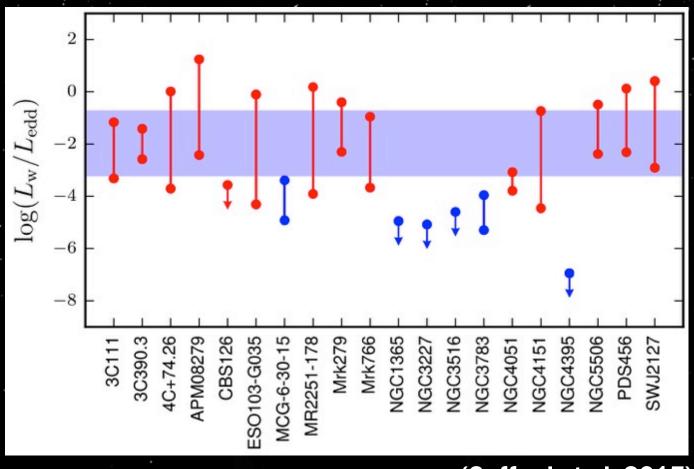


L_K/L_{bol}~0.5-5% (e.g., Di Matteo+ 2005; Hopkins & Elvis 2010; Gaspari+ 2011)

F. Tombesi - AGN Driven Winds, 22-26 May 2017, Technion, Israel

Fombesi et al. (2013)

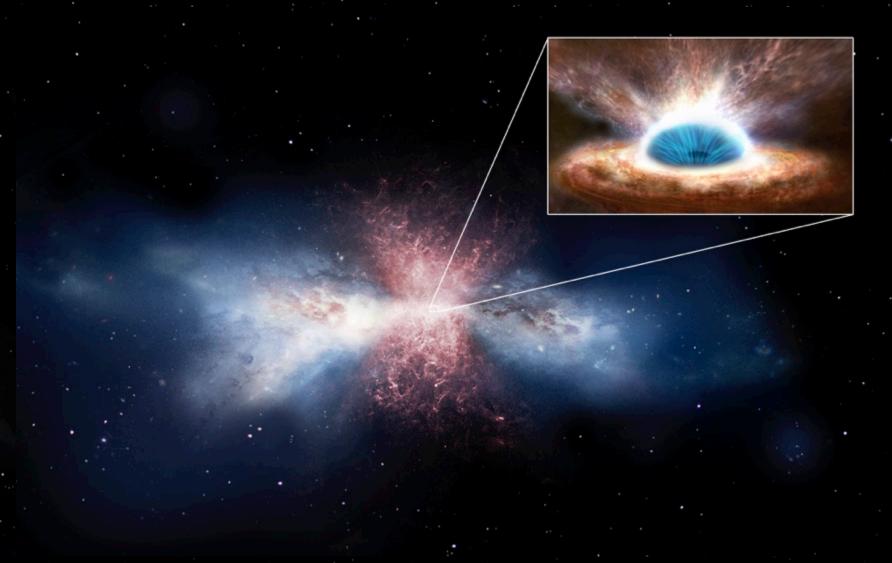
How powerful are ultrafast outflows?



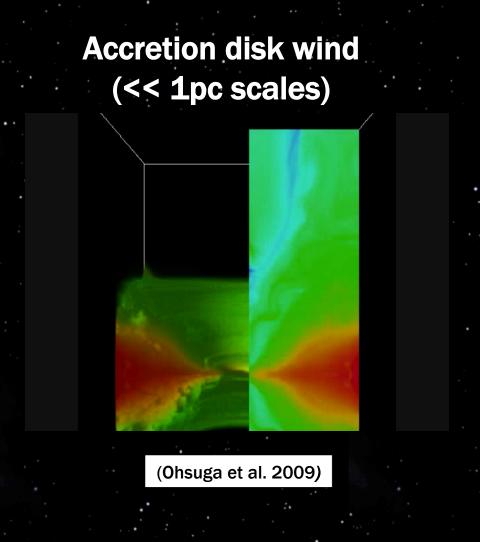
(Gofford et al. 2015)

$$L_W = \frac{1}{2}\dot{M}_{out}v_{out}^2 \sim \Omega N_H R_{in}v_{out}^3$$

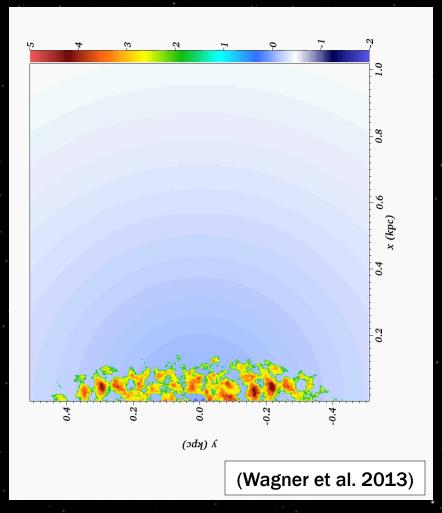
5 - Are there evidences for AGN wind feedback in action?



AGN feedback from accretion disk winds?



Galaxy feedback (kpc scales)



nature

THE INTERNATIONAL WEEKLY JOURNAL OF SCIENCE

GROWING IN THE

Accretion-disk winds drive evolution of supermassive black holes and their galaxies

PAGES 423 & 436

INNOVATION

BIOTECH BOOT CAMP

Finishing school for wannabe entrepreneurs MEDICAL GENETICS

"DON'T EDIT THE GERM LINE"

Heritable gene modification risks may outweigh benefits PAGE 410

NAMOTECHNOLOGY

WATER ENTERS A NEW PHASE

'Square ice' found between the graphene sheets PAGES 417 & 443

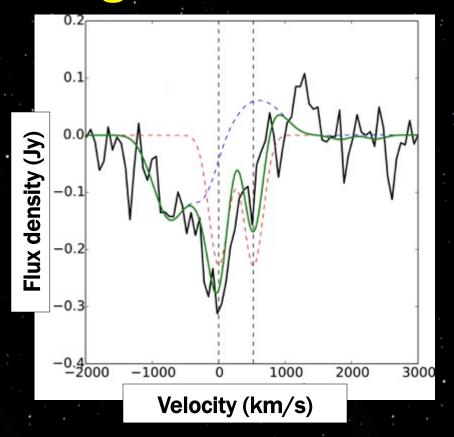
⇒ NATURE.COM/NATURE

26 March 2015



Tombesi et al. (2015)

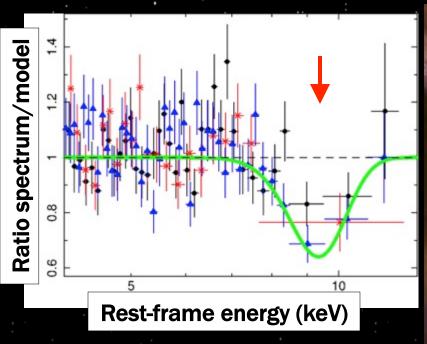
Large-scale molecular outflow with Herschel





- IRAS F11119+3257, ULIRG z=0.189, QSO luminosity 10⁴⁶erg/s
- Herschel spectrum OH 119µm P-Cygni line profile (Veilleux et al. 2013)
- Molecular outflow 1000 km/s, 800 M_{\odot} yr⁻¹ at >300 pc
- Follow-up with ALMA (PI Veilleux). ApJ paper under review

Accretion disk wind detected with Suzaku

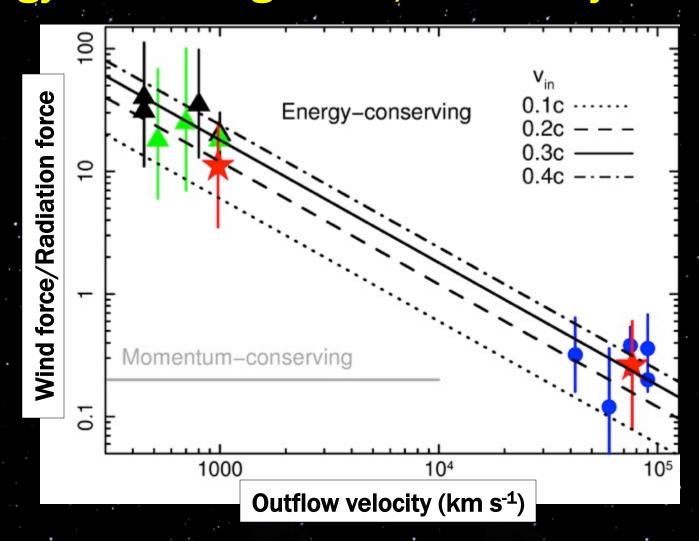




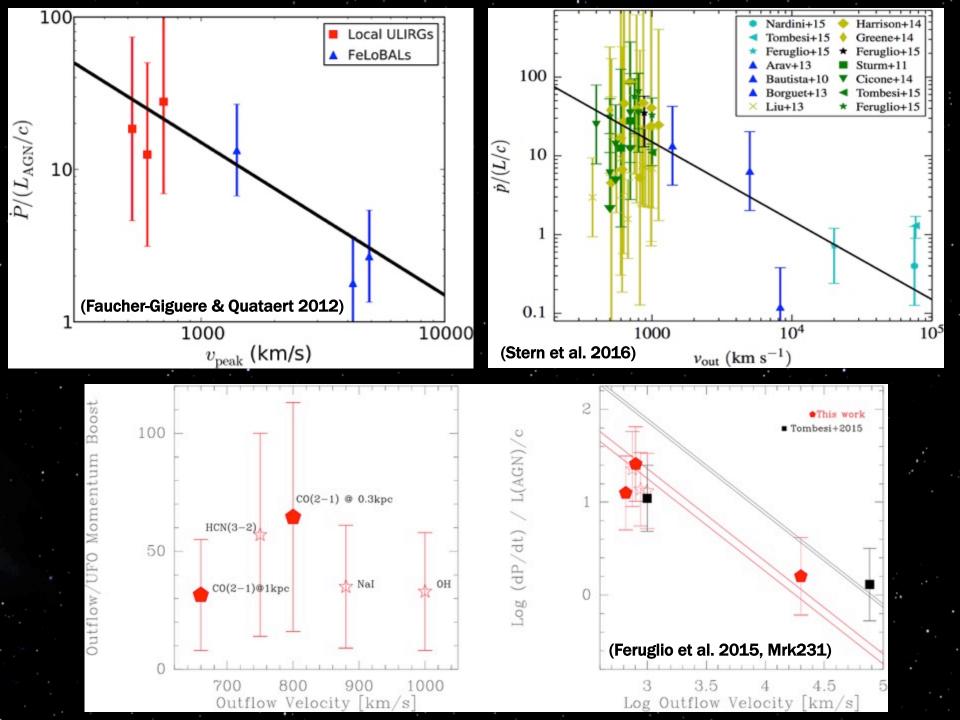
(Image credit ESA)

- Long 250ks Suzaku observation in May 2013
- Detection (6.5sigma) broad absorption line at rest-frame E=9.82keV
- XSTAR fit: v=0.255c, logxi=4.11, Nh=6×10²⁴, covering fraction >0.85
- Follow-up with NuSTAR (PI Tombesi). Paper to be submitted to ApJ soon

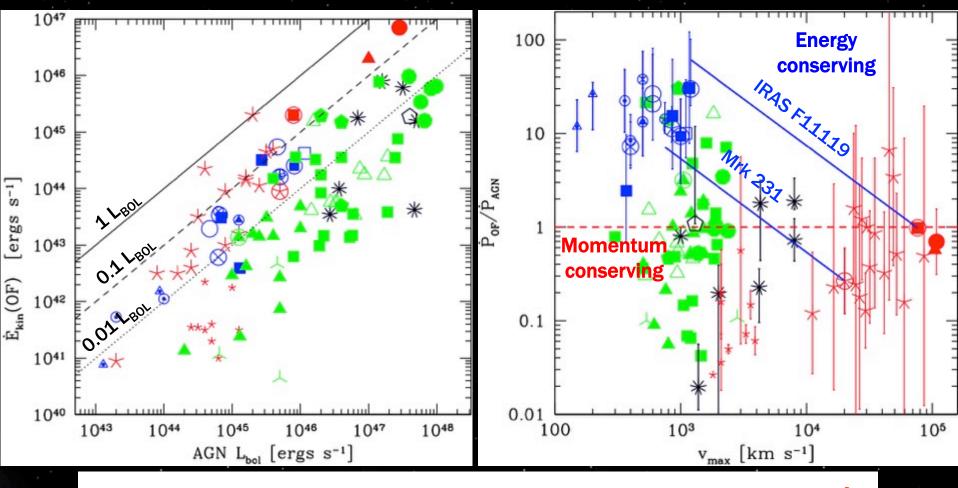
Energy conserving winds, from X-ray to IR/mm



- Energy-conserving flow: $P_{out} = f(v_{in}/v_{out})(L_{AGN}/c)$
- Efficiency energy conservation f=C_{F,OH}/C_{F,X}=0.2



AGN wind scaling relations



Molecular winds Ionized winds BAL winds X-ray winds (WA*, UFO*)

(Fiore et al. 2017)

Galaxy-scale feedback from AGN winds

(credit NASA/GSFC press release)

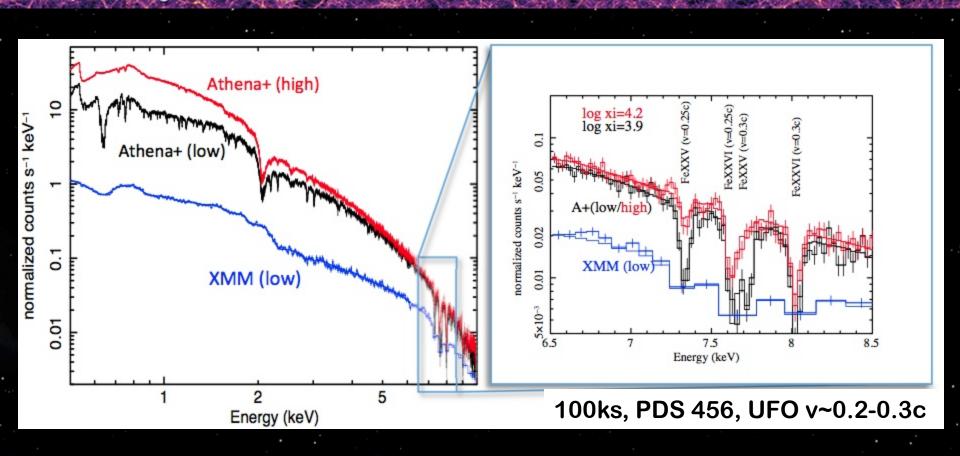
ATHENA

THE ASTROPHYSICS OF THE HOT AND ENERGETIC UNIVERSE

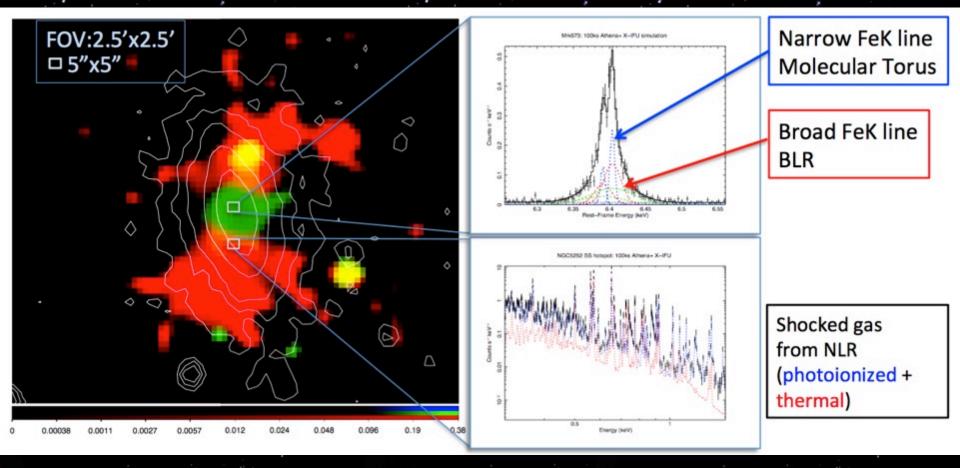


HOW DO BLACK HOLES GROW AND SHAPE THE UNIVERSE?

Europe's next generation X-RAY OBSERVATORY



Spectral-imaging with X-IFU: Seyfert 2 NGC 5252



(Cappi et al. 2013, adapted from Dadina et al. 2010)

<u>Multiwavelength campaigns with ALMA, VLT, E-ELT</u>: hot/cold gas mixing, AGN wind power, wind bubbles, shocks, photoionization cones, outflow morphology, AGN/star formation comparison, ... see also AXIS and Lynx mission proposals to NASA!

